

# Dynamics of Staircase Formation



Reduced Model of Beta Plane Turbulence

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## **Abstract / Introduction/Objectives**

**Abstract** A two-field model for staircase dynamics relevant to both beta-plane geostrophic and drift-wave plasma turbulence is studied numerically and analytically. The model evolves an averaged potential vorticity (PV) whose flux is both driven by, and regulates, an enstrophy field,  $\varepsilon$ . The model's closure uses a mixing length concept. Its link with bistability, vital to staircase generation, is analysed and verified by integrating the equations numerically.

Introduction The turbulent transport and structure formation phenomenon known as a 'staircase', originally introduced in [2] manifests itself as follows:

- stably stratified density profile in the ocean occasionally reorganizes into layers separated by thin interfaces
- positive feedback provided by a profile rippling instability is equivalent to a 'negative diffusivity' that enhances the profile corrugation instead of smoothing it
- negative diffusion corresponds to a descending branch of an "S-curve" in the flux - gradient relation, i.e. a range of  $\nabla n$  for which  $\delta\Gamma/\delta(-\nabla n) < 0$

#### Objectives

- 1. identification of conditions and the parameter space for staircase formation
- 2. demonstration of staircase persistence by direct numerical integration of the model equations
- 3. finding exact analytic steady state solutions

- density gradient flattens in the layers and steepens in the interfaces  $\rightarrow$ 'staircase'
- pre-existing turbulent transport is supported by, and regulates, the gradient
- feedback loop drives the transport supporting turbulence out of the regions with steeper profiles into adjacent regions with the flatter ones, thus settling at a *bistable* equilibrium

(2)

- and exploiting them for code verification
- 4. elucidation of staircase dynamics, long time evolution, merger events and the role of domain boundaries

Model:

$$Q_t = \partial_y \frac{\varepsilon^{1/2}}{\left(1 + Q_y^2/\varepsilon\right)^2} Q_y + DQ_y$$

$$\varepsilon_{t} = \partial_{y} \frac{\varepsilon^{1/2}}{\left(1 + Q_{y}^{2}/\varepsilon\right)^{2}} \varepsilon_{y} + D\varepsilon_{yy} + \frac{\varepsilon^{1/2}}{\left(1 + Q_{y}^{2}/\varepsilon\right)^{2}} Q_{y}^{2} - \frac{\varepsilon^{3/2}}{\varepsilon_{0}} + \gamma \sqrt{\varepsilon}$$

**Formulation** Consider potential vorticity (PV), q, of a geostrophic fluid, e.g., on a rapidly rotating planet. It consists of the planetary vorticity (on  $\beta$ -plane) and fluid vorticity:

$$q = \beta y + \Delta \psi$$

where  $\psi$  is the stream function, and y is a latitudinal coordinate. Equation for q:

$$\frac{\partial q}{\partial t} - \nabla \psi \times \nabla q = \nu \Delta \psi + f \tag{1}$$

Decompose q and  $\psi$  into a mean and fluctuating parts

$$q = \langle q(y,t) \rangle + \tilde{q}(x,y,t)$$

with  $\tilde{q} = \Delta \tilde{\psi}$ . Separate the *x*-averaged component  $Q \equiv \langle q \rangle$  from fluctuating part squared (enstrophy),  $\varepsilon = \langle \tilde{q}^2 \rangle / 2$ . The closure

problem for  $\langle \nabla \tilde{\psi} \times \nabla \Delta \tilde{\psi} \rangle$  arises. For fluctuations statistically homogeneous in *x*-direction the *x*-averaged PV flux  $\Gamma_q$  is:

$$-\frac{\partial\Gamma_q}{\partial y} \equiv \langle\nabla\tilde{\psi}\times\nabla\Delta\tilde{\psi}\rangle = \frac{\partial^2}{\partial y^2} \left\langle\frac{\partial\tilde{\psi}}{\partial x}\frac{\partial\tilde{\psi}}{\partial y}\right\rangle.$$

Next, we apply a Fickian *Ansatz*:  $\Gamma_q = -D_e \partial Q / \partial y$ , where  $D_e(\varepsilon, Q_y)$  is the PV diffusivity. This is assumed to follow a mixing-length hypothesis,  $D_e \sim l |\nabla \tilde{\psi}|$ , where  $l(\varepsilon, Q_y)$  is the mixing length, introduced phenomenologically as [1]:

$$\frac{1}{2} = \frac{1}{l_0^2} + \frac{1}{l_R^2}.$$

Here,  $l_0$  is a fixed contribution to the mixing length l that characterizes the turbulence, e.g., the stirring scale.  $l_R$  is the Rhines scale at which dissipation of  $\varepsilon$  balances its production, so  $l_R =$ 

 $l_R(\varepsilon, Q_y)$ . In turbulent cascades where wave form of energy coexists with turbulent eddies, the Rhines scale is where these two intersect, i.e., where  $k\tilde{v} \sim \omega_k$  [3]. When the turbulent energy inverse cascade reaches this scale, it is intercepted and transported further by waves both in wave-number and configuration space. The only dimensionless combination of the variables entering eq.(2) is  $l_0^2 Q_y^2 / \varepsilon$ . So, we may generalize the relation in eq.(2) and write  $l_0/l = \left(1 + l_0^2 Q_y^2 / \varepsilon\right)^{\kappa}$ . We choose  $\kappa = 2$ . Replacing the eddy velocity in the Fick's law by  $l_0\sqrt{\varepsilon}$  and measuring y in units of  $l_0$ , we can write the averaged eq.(1) for Q as shown above with added (small) constant diffusivity D. Applying similar arguments to the turbulent part of PV,  $\varepsilon$ , and adding the terms responsible for its production, damping and unstable growth, we obtain the above evolution equation for the enstrophy  $\varepsilon$ .

## **Staircase Prerequisites/Formation/Merger**

• SC result from the loss of stability of a ground state solution for Q and  $\epsilon$ , characterized by the constant values  $\epsilon = \epsilon_B$  and  $Q_y = Q'_B$  that annihilate the enstrophy production-dissipation term:





• stationary SC structure is a quasi-periodic sequence of regions with alternating upper and lower stable  $\varepsilon$  values

• time-asymptotically, this solution can be calculated analytically





#### the central two steps merge into a bigger step



• flux remains constant when no mergers occur

 $\left[\varepsilon^{1/2}\left(1+Q_y^2/\varepsilon\right)^{-2}+D\right]Q_y \equiv b = const$ 

• the flux builds up in two phases (slow and fast) before it drops abruptly to its averaged value after the merger

- the first phase is an initial growth that lasts to about  $t \approx 0.065$ . The flux increase remains relatively small, < 0.01
- the second phase is explosive and can be accurately fit by the following function,

Numerical solution in long-time asymptotic regime, shown with the solid line. Exact analytic solution represented by the two branches shown with red and green squares ( $\psi = Q_y / \sqrt{\varepsilon}$ )



• quasi-stationary SC forms quickly ( $t \ll 1$ ) with *n* steps separated by shear layers with steep gradient of the mean vorticity  $Q_y$  and suppressed enstrophy level,  $\varepsilon$ 

• number *n* is determined by the maximum growth rate

over a longer time (but still ≪ 1), most of *n* steps merge with their neighbours and the total number of steps becomes ≈ *n*/2. After this initial phase the staircase persists for a much longer time



*Q*-flux grows rapidly, and strongly deviates from its globally constant value precisely at the merger locations
shown is a sequence of mergers of 12 initial steps. They proceed symmetrically from the boundaries towards the centre
process continues until the mergers converge at the centre and

#### $F = F_0 + B / (t_0 - t)^{\alpha}$

with  $t_0 \approx 0.0863$ ,  $B \approx 0.000806$ ,  $\alpha \approx 0.879$ , and a residual flux  $F_0 \approx -0.0171$ .

## References

 [1] BALMFORTH, N. J., SMITH, S. G. L. & YOUNG, W. R. 1998 Dynamics of interfaces and layers in a stratified turbulent fluid. *Journal of Fluid Mechanics* 355, 329–358.

[2] PHILLIPS, O. M. 1972 Turbulence in a strongly stratified fluid
– is it unstable? *Deep Sea Research and Oceanographic Abstracts* 19, 79–81.

[3] RHINES, P. B. 1975 Waves and turbulence on a beta-plane. *Journal of Fluid Mechanics* **69**, 417–443.