

# Turbulence and Transport in Elastic Media: A New Look at Classic Themes

Xiang Fan<sup>1</sup>, P H Diamond<sup>1</sup>, Luis Chacon<sup>2</sup>

<sup>1</sup> University of California, San Diego

<sup>2</sup> Los Alamos National Laboratory

Sherwood 2018

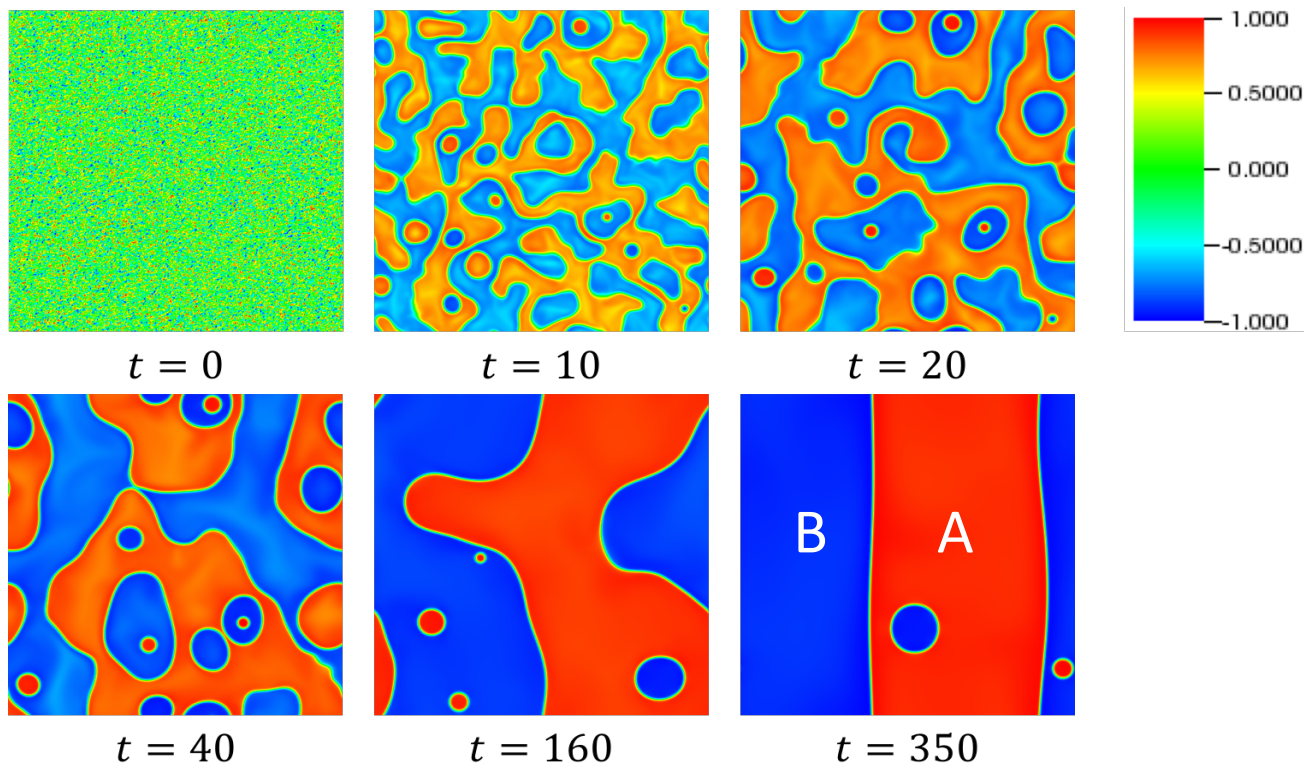
This research was supported by the U.S. Department of Energy, Office of Science, Office of Fusion Energy Sciences, under Award Number DE-FG02-04ER54738 and CMTFO.

# “Tour Guide”

- This is NOT a traditional study of plasma physics.
- It is about a *new* system that is related to systems you are familiar with in plasma physics
- There are many similarities, but some important differences. Watch for these!
- We studied the fundamental physics of cascades and self-organization in this system and in MHD
- It provides a new look at classic themes in plasma physics.

# Elastic Media? -- What Is the CHNS System?

- Elastic media – Fluid with internal DoFs → “springiness”
- The Cahn-Hilliard Navier-Stokes (CHNS) system describes ***phase separation*** for binary fluid (i.e. ***Spinodal Decomposition***)



Miscible phase  
→ Immiscible phase

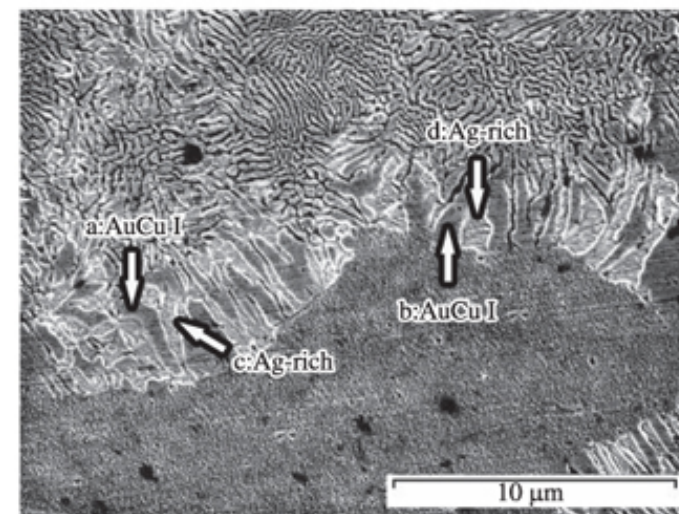


Figure 5. FE-SEM micrograph of specimen aged at 400 °C for 5000 minutes.

# Elastic Media? -- What Is the CHNS System?

- How to describe the system: the concentration field
- $\psi(\vec{r}, t) \stackrel{\text{def}}{=} [\rho_A(\vec{r}, t) - \rho_B(\vec{r}, t)]/\rho$  : scalar field
- $\psi \in [-1, 1]$
- CHNS equations:

$$\partial_t \psi + \vec{v} \cdot \nabla \psi = D \nabla^2 (-\psi + \psi^3 - \xi^2 \nabla^2 \psi)$$

$$\partial_t \omega + \vec{v} \cdot \nabla \omega = \frac{\xi^2}{\rho} \vec{B}_\psi \cdot \nabla \nabla^2 \psi + \nu \nabla^2 \omega$$

# Outline

- A Brief Derivation of the CHNS Model
- 2D CHNS and 2D MHD
- Linear Wave
- Ideal Quadratic Conserved Quantities
- Scales, Ranges, Trends
- Cascades
- Power Laws
- Transport
- Conclusions

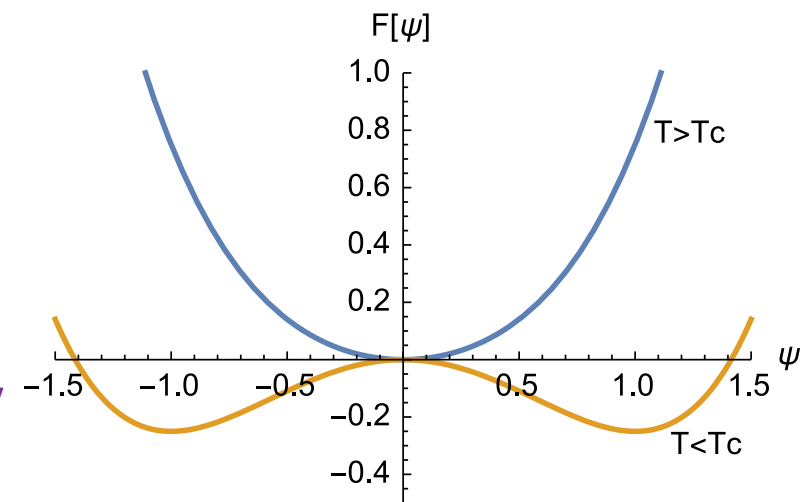
# A Brief Derivation of the CHNS Model

- Second order phase transition  $\rightarrow$  Landau Theory.
- Order parameter:  $\psi(\vec{r}, t) \stackrel{\text{def}}{=} [\rho_A(\vec{r}, t) - \rho_B(\vec{r}, t)]/\rho$
- Free energy:

$$F(\psi) = \int d\vec{r} \left( \underbrace{\frac{1}{2} C_1 \psi^2 + \frac{1}{4} C_2 \psi^4}_{\text{Phase Transition}} + \underbrace{\frac{\xi^2}{2} |\nabla \psi|^2}_{\text{Gradient Penalty}} \right)$$

- $C_1(T), C_2(T)$ .
- Isothermal  $T < T_C$ . Set  $C_2 = -C_1 = 1$ :

$$F(\psi) = \int d\vec{r} \left( -\frac{1}{2} \psi^2 + \frac{1}{4} \psi^4 + \frac{\xi^2}{2} |\nabla \psi|^2 \right)$$



# A Brief Derivation of the CHNS Model

- Continuity equation:  $\frac{d\psi}{dt} + \nabla \cdot \vec{J} = 0$ . Fick's Law:  $\vec{J} = -D\nabla\mu$ .

- Chemical potential:  $\mu = \frac{\delta F(\psi)}{\delta\psi} = -\psi + \psi^3 - \xi^2 \nabla^2 \psi$ .

- Combining above  $\rightarrow$  Cahn Hilliard equation:

$$\frac{d\psi}{dt} = D\nabla^2 \mu = D\nabla^2 (-\psi + \psi^3 - \xi^2 \nabla^2 \psi)$$

- $d_t = \partial_t + \vec{v} \cdot \nabla$ . Surface tension: force in Navier-Stokes equation:

$$\partial_t \vec{v} + \vec{v} \cdot \nabla \vec{v} = -\frac{\nabla p}{\rho} - \psi \nabla \mu + \nu \nabla^2 \vec{v}$$

- For incompressible fluid,  $\nabla \cdot \vec{v} = 0$ .

# 2D CHNS and 2D MHD

- 2D CHNS Equations:

$$\partial_t \psi + \vec{v} \cdot \nabla \psi = D \nabla^2 (-\psi + \psi^3 - \xi^2 \nabla^2 \psi)$$

$$\partial_t \omega + \vec{v} \cdot \nabla \omega = \frac{\xi^2}{\rho} \vec{B}_\psi \cdot \nabla \nabla^2 \psi + \nu \nabla^2 \omega$$

$-\psi$ : Negative diffusion term

$\psi^3$ : Self nonlinear term

$-\xi^2 \nabla^2 \psi$ : Hyper-diffusion term

With  $\vec{v} = \hat{z} \times \nabla \phi$ ,  $\omega = \nabla^2 \phi$ ,  $\vec{B}_\psi = \hat{z} \times \nabla \psi$ ,  $j_\psi = \xi^2 \nabla^2 \psi$ .

- 2D MHD Equations:

$$\partial_t A + \vec{v} \cdot \nabla A = \eta \nabla^2 A$$

$$\partial_t \omega + \vec{v} \cdot \nabla \omega = \frac{1}{\mu_0 \rho} \vec{B} \cdot \nabla \nabla^2 A + \nu \nabla^2 \omega$$

$A$ : Simple diffusion term

With  $\vec{v} = \hat{z} \times \nabla \phi$ ,  $\omega = \nabla^2 \phi$ ,  $\vec{B} = \hat{z} \times \nabla A$ ,  $j = \frac{1}{\mu_0} \nabla^2 A$ .

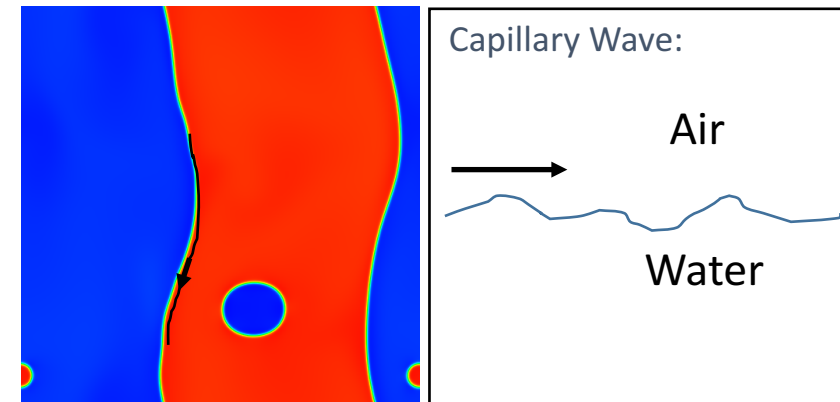
	2D MHD	2D CHNS
Magnetic Potential	$A$	$\psi$
Magnetic Field	$\mathbf{B}$	$\mathbf{B}_\psi$
Current	$j$	$j_\psi$
Diffusivity	$\eta$	$D$
Interaction strength	$\frac{1}{\mu_0}$	$\xi^2$



# Linear Wave

- CHNS supports linear “elastic” wave:

$$\omega(k) = \pm \sqrt{\frac{\xi^2}{\rho} |\vec{k} \times \vec{B}_{\psi_0}|} - \frac{1}{2} i(CD + \nu)k^2$$



Where  $C \equiv [-1 - 6\psi_0 \nabla^2 \psi_0 / k^2 - 6(\nabla \psi_0)^2 / k^2 - 6\psi_0 \nabla \psi_0 \cdot i\mathbf{k} / k^2 + 3\psi_0^2 + \xi^2 k^2]$

- Akin to capillary wave at phase interface. Propagates ***only*** along the interface of the two fluids, where  $|\vec{B}_{\psi}| = |\nabla \psi| \neq 0$ .
- Analogue of Alfvén wave.
- Important differences:
  - $\vec{B}_{\psi}$  in CHNS is large only in the interfacial regions.
  - Elastic wave activity does not fill space.

# Ideal Quadratic Conserved Quantities

## • 2D MHD

### 1. Energy

$$E = E^K + E^B = \int \left( \frac{v^2}{2} + \frac{B^2}{2\mu_0} \right) d^2x$$

### 2. Mean Square Magnetic Potential

$$H^A = \int A^2 d^2x$$

### 3. Cross Helicity

$$H^C = \int \vec{v} \cdot \vec{B} d^2x$$

## • 2D CHNS

### 1. Energy

$$E = E^K + E^B = \int \left( \frac{v^2}{2} + \frac{\xi^2 B_\psi^2}{2} \right) d^2x$$

### 2. Mean Square Concentration

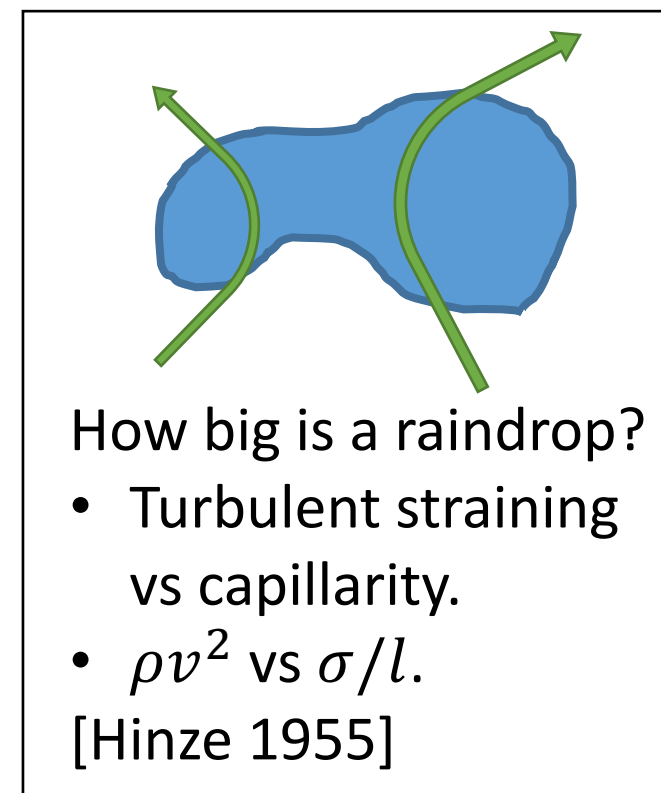
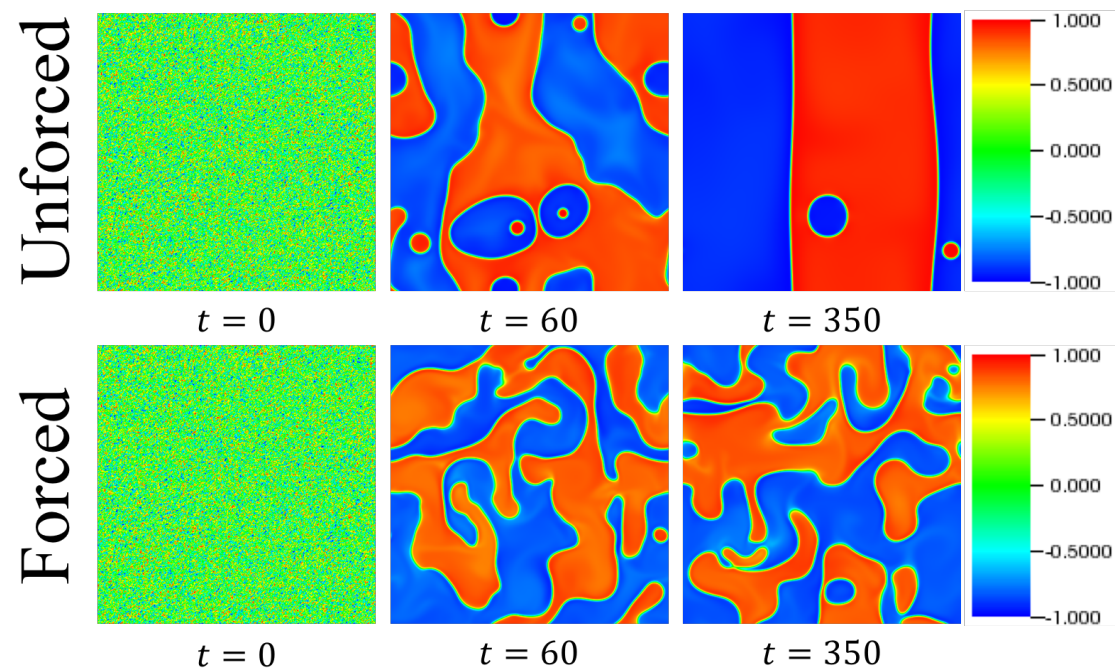
$$H^\psi = \int \psi^2 d^2x$$

### 3. Cross Helicity

$$H^C = \int \vec{v} \cdot \vec{B}_\psi d^2x$$

Dual cascade expected!

# Scales, Ranges, Trends

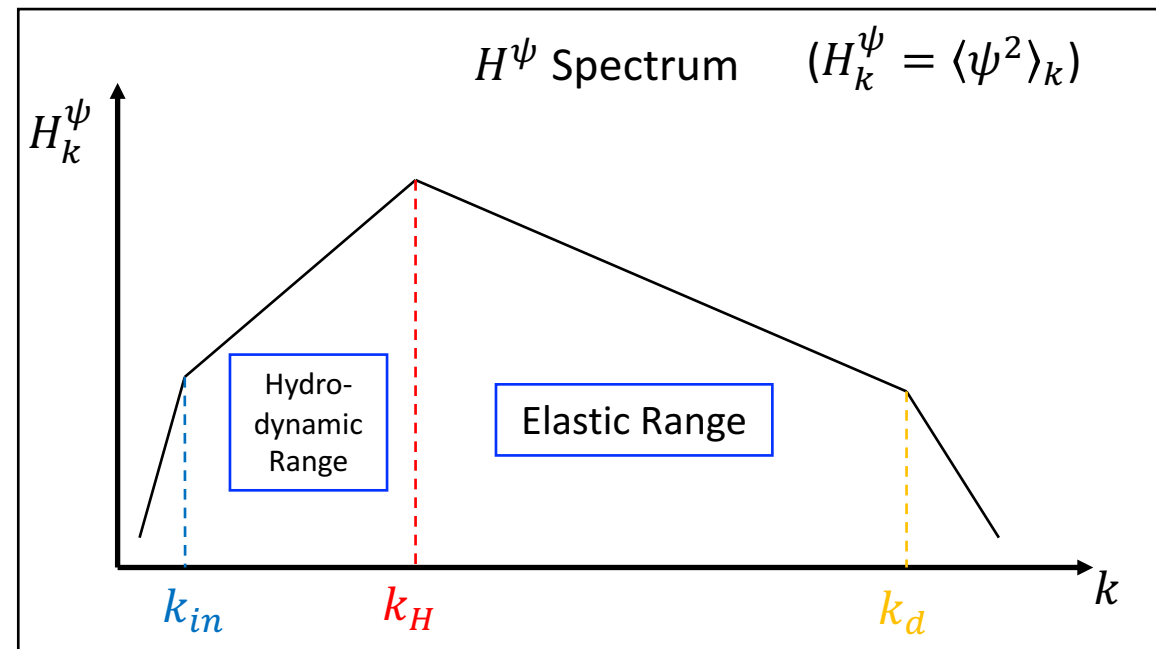


- Fluid forcing  $\rightarrow$  Fluid straining vs Blob coalescence
- Scale where turbulent straining  $\sim$  elastic restoring force (due surface tension): Hinze Scale

$$L_H \sim \left(\frac{\rho}{\xi}\right)^{-1/3} \epsilon_{\Omega}^{-2/9}$$

# Scales, Ranges, Trends

- Elastic range:  $L_H < l < L_d$ : where elastic effects matter.
- $L_H/L_d \sim \left(\frac{\rho}{\xi}\right)^{-1/3} \nu^{-1/2} \epsilon_\Omega^{-1/18} \rightarrow$  Extent of the elastic range
- $L_H \gg L_d$  required for large elastic range  $\rightarrow$  case of interest

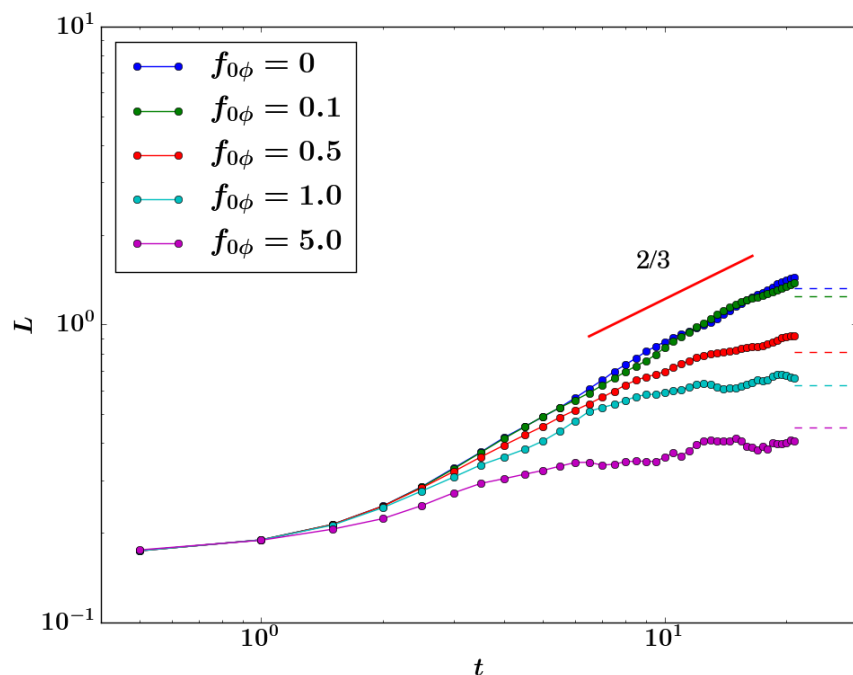
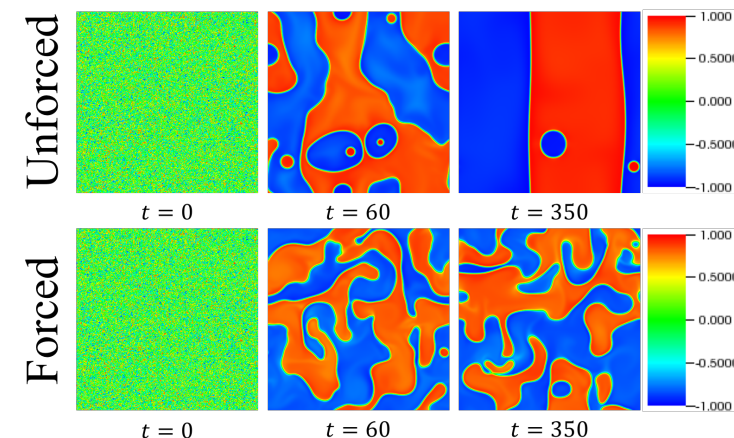


# Scales, Ranges, Trends

- Key elastic range physics: **Blob coalescence**
- Unforced case:  $L(t) \sim t^{2/3}$ .

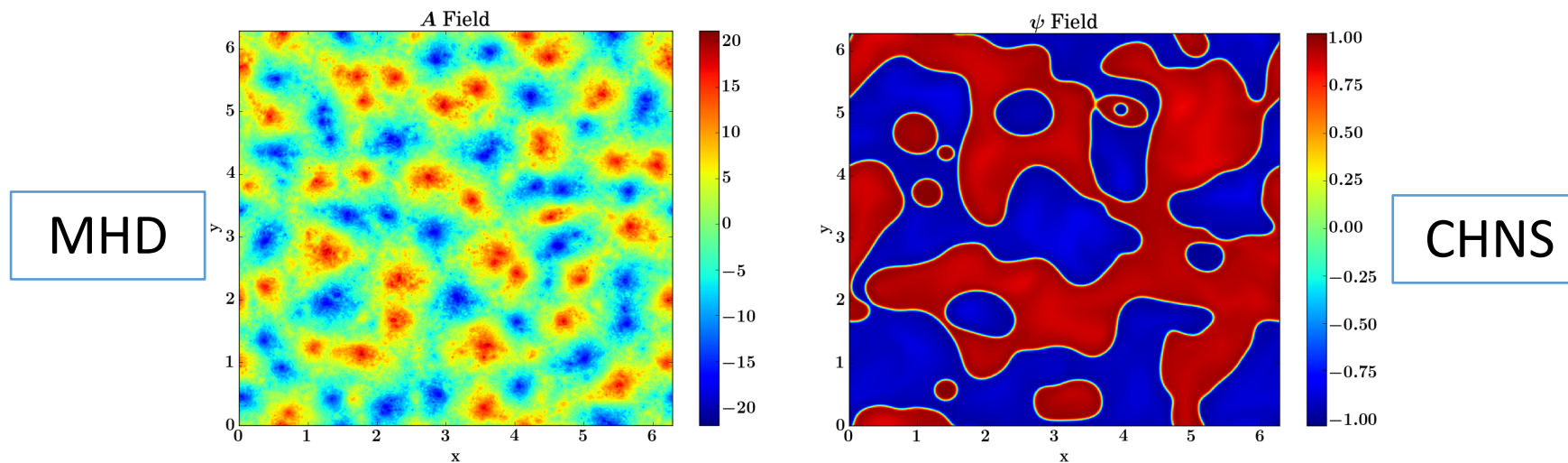
$$\text{(Derivation: } \vec{v} \cdot \nabla \vec{v} \sim \frac{\xi^2}{\rho} \nabla^2 \psi \nabla \psi \Rightarrow \frac{\dot{L}^2}{L} \sim \frac{\sigma}{\rho} \frac{1}{L^2} \text{)}$$

- Forced case: blob coalescence arrested at Hinze scale  $L_H$ .



- $L(t) \sim t^{2/3}$  recovered
- Blob growth arrest observed
- Blob growth saturation scale tracks Hinze scale (dashed lines)

# Cascades



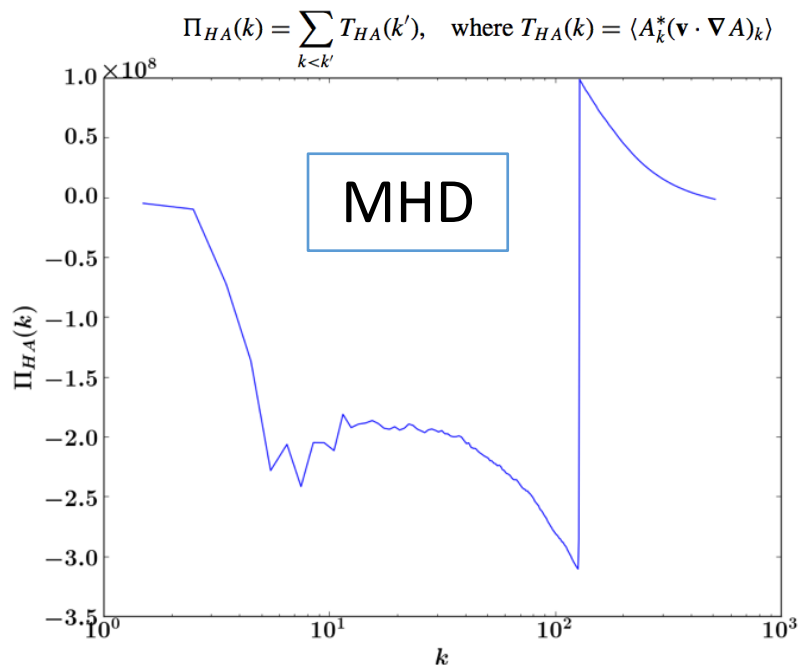
- Blob coalescence in the elastic range of CHNS is analogous to flux coalescence in MHD.
- Suggests *inverse cascade* of  $\langle \psi^2 \rangle$  in CHNS.
- Supported by the statistical mechanics studies (absolute equilibrium distributions).

# Cascades

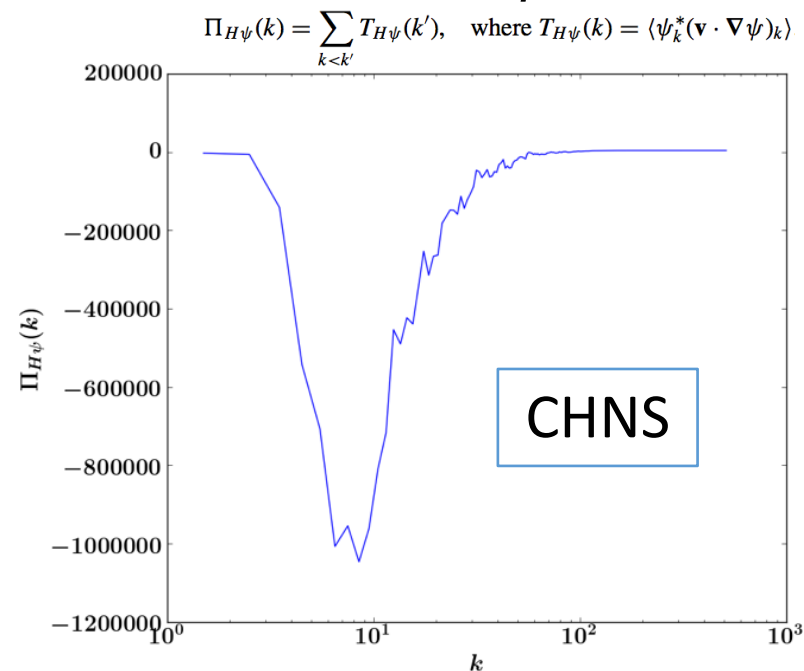
- So, dual cascade:
  - *Inverse* cascade of  $\langle \psi^2 \rangle$
  - *Forward* cascade of  $E$
- Inverse cascade of  $\langle \psi^2 \rangle$  is formal expression of blob coalescence process  $\rightarrow$  generate larger scale structures till limited by straining
- Forward cascade of  $E$  as usual, as elastic force breaks enstrophy conservation

# Cascades

- Spectral flux of  $\langle A^2 \rangle$ :



- Spectral flux of  $\langle \psi^2 \rangle$ :

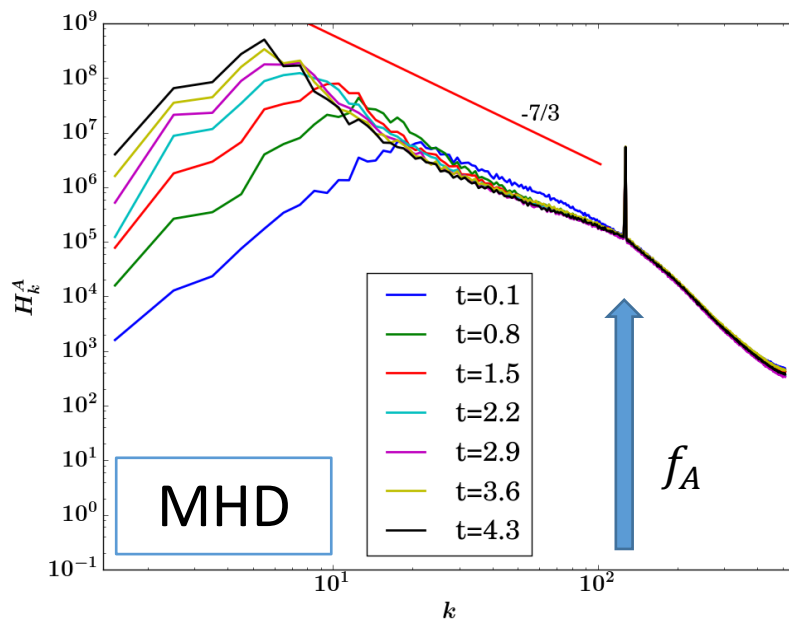


- MHD: weak small scale forcing on  $A$  drives inverse cascade
- CHNS:  $\psi$  is unforced  $\rightarrow$  aggregates naturally
- Both fluxes **negative**  $\rightarrow$  **inverse** cascades

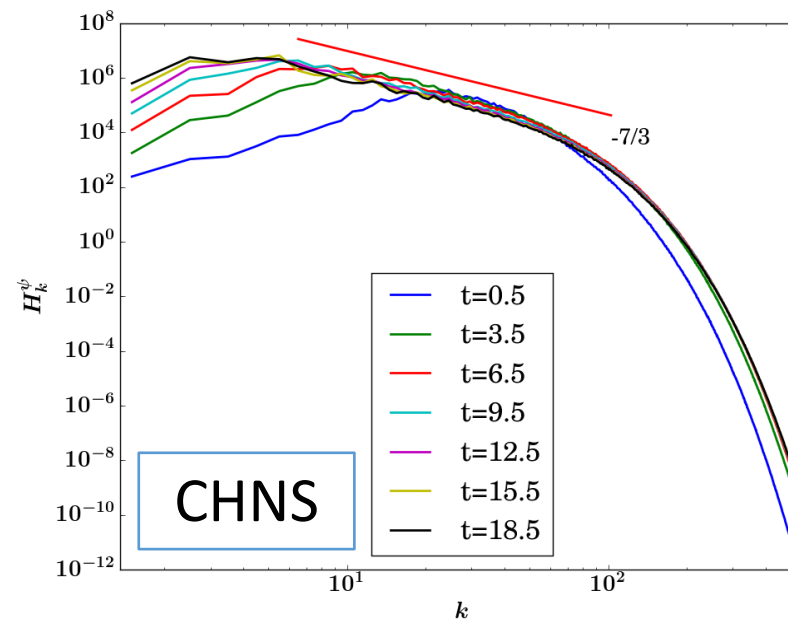


# Power Laws

- $\langle A^2 \rangle$  spectrum:



- $\langle \psi^2 \rangle$  spectrum:



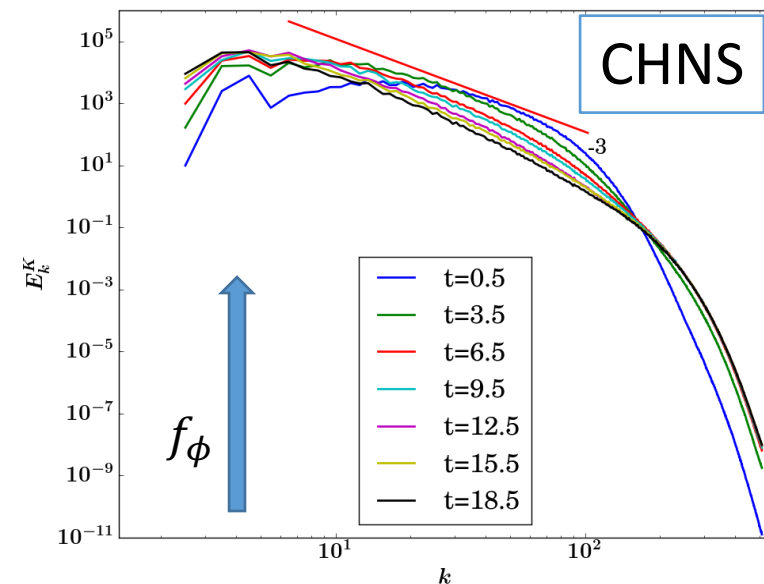
- Both systems exhibit  $k^{-7/3}$  spectra.
- Inverse cascade of  $\langle \psi^2 \rangle$  exhibits same power law scaling, so long as  $L_H \gg L_d$ , maintaining elastic range: Robust process.

# Power Laws

- Derivation of -7/3 power law:
- For MHD, key assumptions:
  - Alfvénic equipartition ( $\rho \langle v^2 \rangle \sim \frac{1}{\mu_0} \langle B^2 \rangle$ )
  - Constant mean square magnetic potential dissipation rate  $\epsilon_{HA}$ , so
 
$$\epsilon_{HA} \sim \frac{H^A}{\tau} \sim (H_k^A)^{\frac{3}{2}} k^{\frac{7}{2}}.$$
- Similarly, assume the following for CHNS:
  - Elastic equipartition ( $\rho \langle v^2 \rangle \sim \xi^2 \langle B_\psi^2 \rangle$ )
  - Constant mean square magnetic potential dissipation rate  $\epsilon_{H\psi}$ , so
 
$$\epsilon_{H\psi} \sim \frac{H^\psi}{\tau} \sim (H_k^\psi)^{\frac{3}{2}} k^{\frac{7}{2}}.$$

# More Power Laws

- Kinetic energy spectrum (**Surprise!**):
- 2D CHNS:  $E_k^K \sim k^{-3}$ ;
- 2D MHD:  $E_k^K \sim k^{-3/2}$ .
- The -3 power law:
  - Closer to enstrophy cascade range scaling, in 2D Hydro turbulence.
  - Remarkable departure from expected -3/2 for MHD. **Why?**
- Why does CHNS  $\leftrightarrow$  MHD correspondence hold well for  $\langle \psi^2 \rangle_k \sim \langle A^2 \rangle_k \sim k^{-7/3}$ , yet break down drastically for energy?
- **What physics** underpins this surprise?

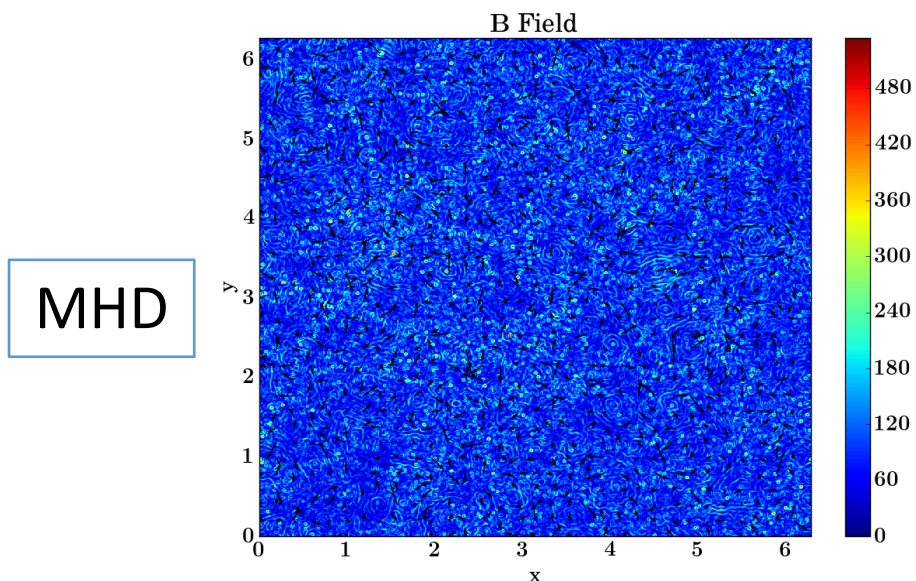


# Interface Packing Matters!

- Need to understand *differences*, as well as similarities, between CHNS and MHD problems.

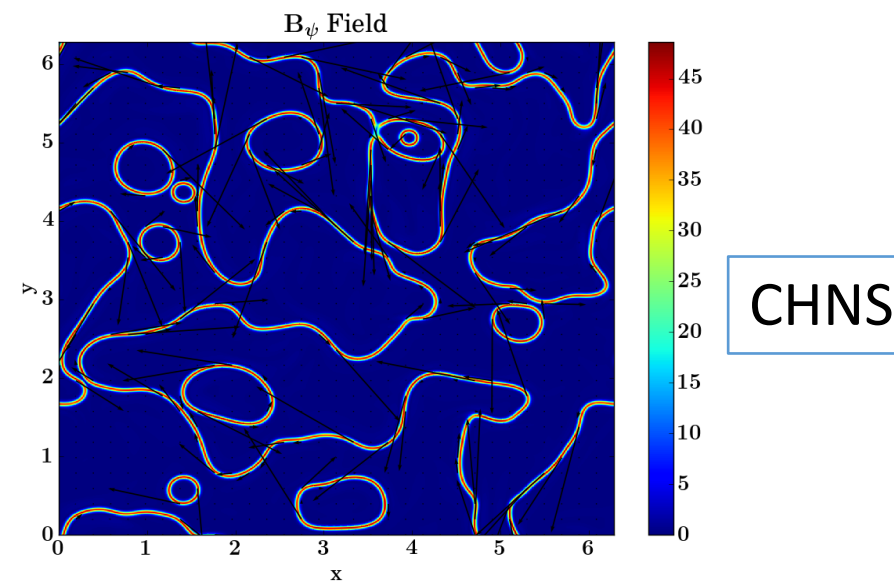
## 2D MHD:

- Fields pervade system.



## 2D CHNS:

- Elastic back-reaction is limited to regions of density contrast i.e.  $|\vec{B}_\psi| = |\nabla\psi| \neq 0$ .
- As blobs coalesce, interfacial region diminished. 'Active region' of elasticity decays.



# Interface Packing Matters!

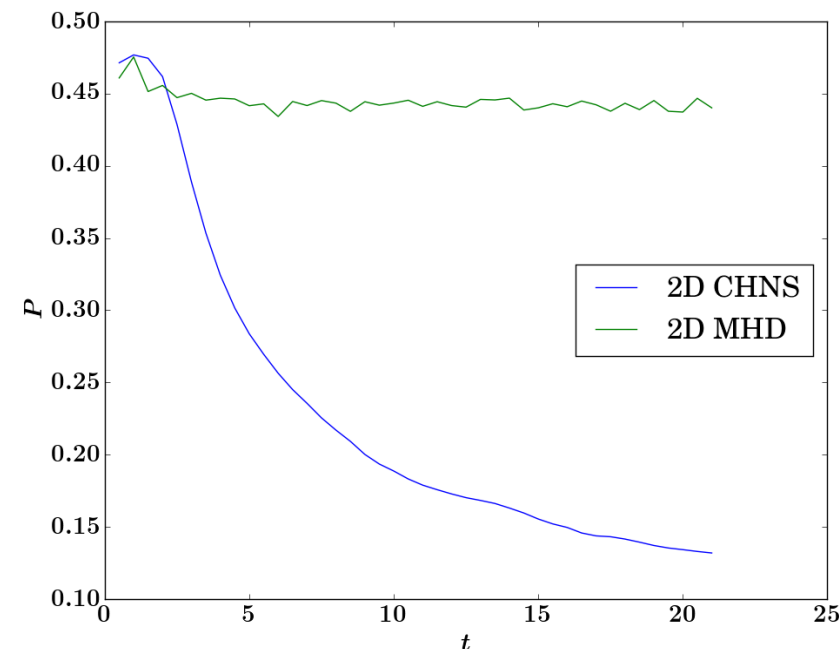
- Define the interface packing fraction  $P$ :

$$P = \frac{\text{\# of grid points where } |\vec{B}_\psi| > B_\psi^{rms}}{\text{\# of total grid points}}$$

➤  $P$  for CHNS decays;

➤  $P$  for MHD stationary!

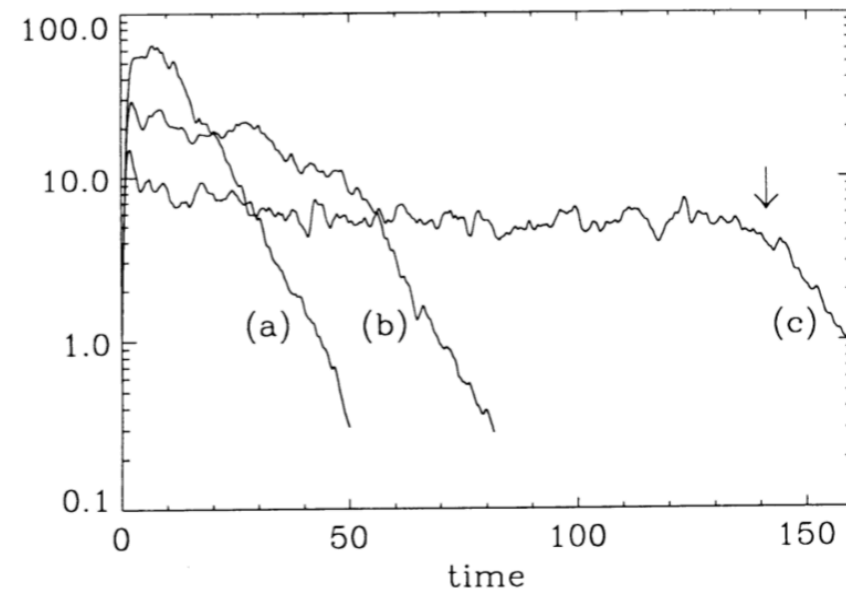
- $\partial_t \omega + \vec{v} \cdot \nabla \omega = \frac{\xi^2}{\rho} \vec{B}_\psi \cdot \nabla \nabla^2 \psi + \nu \nabla^2 \omega$ : small  $P \rightarrow$  local back reaction is weak.
- Weak back reaction  $\rightarrow$  reduce to 2D hydro



# Transport: Something Old

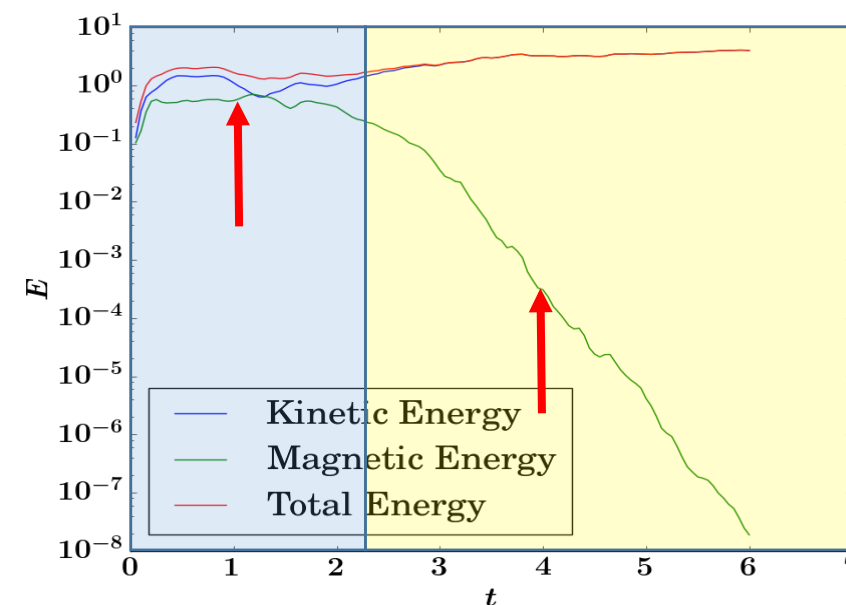
- $M^2 = \langle \tilde{v}^2 \rangle / v_{A0}^2$
- Higher  $v_{A0}^2 / \langle \tilde{v}^2 \rangle \rightarrow$  lower  $D_T \rightarrow$  longer  $E_m$  persistence
- Ultimately  $\eta$  asserts itself

- Blue:  $\langle B \rangle$  sufficient for suppression
- Yellow: Ohmic decay phase



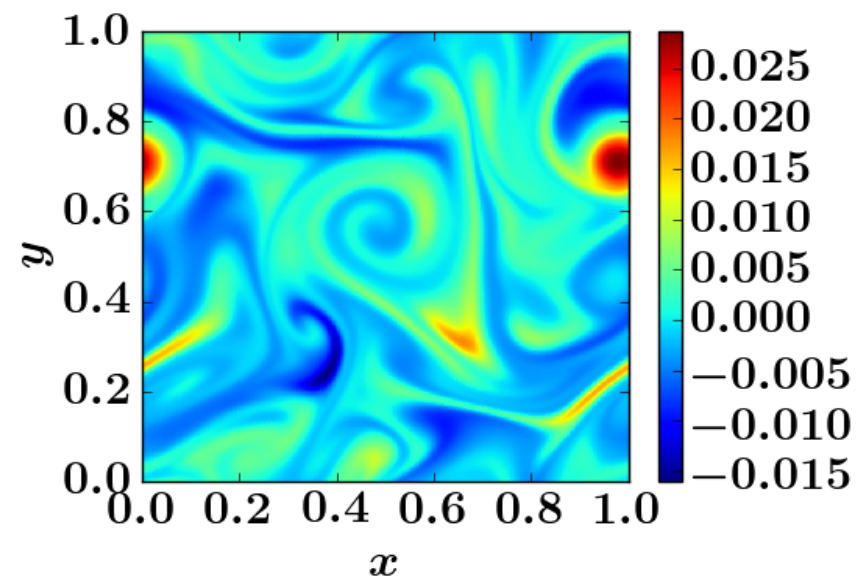
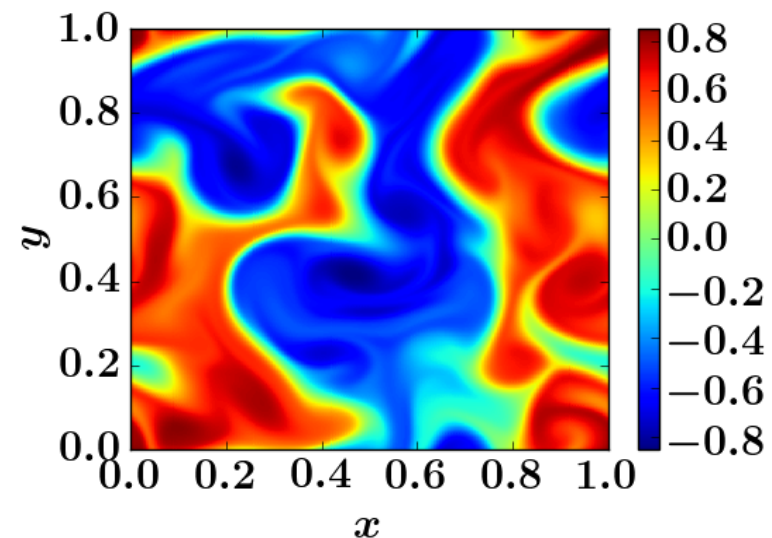
[Cattaneo and Vainshtein '91]

FIG. 3.—Magnetic energy density. Time histories of the total magnetic energy (normalized). The values of  $M^2$  are  $\infty$  for (a), 100 for (b), and 30 for (c).



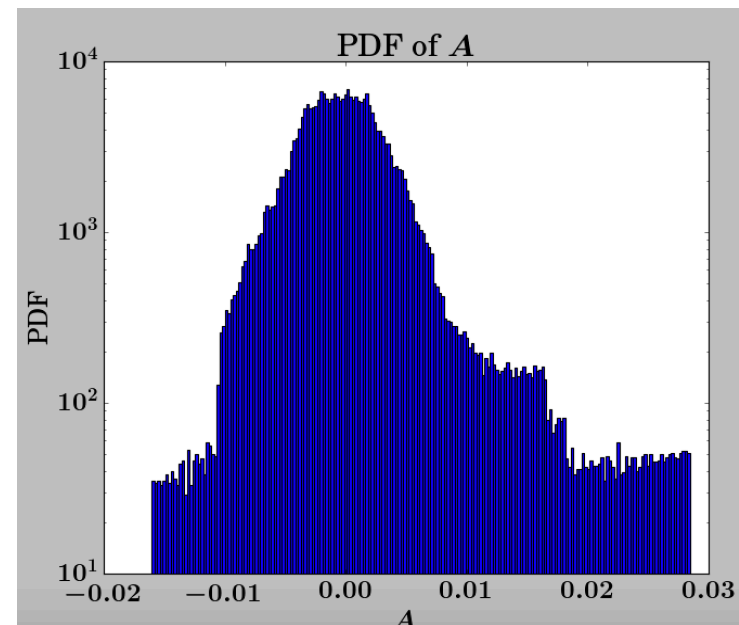
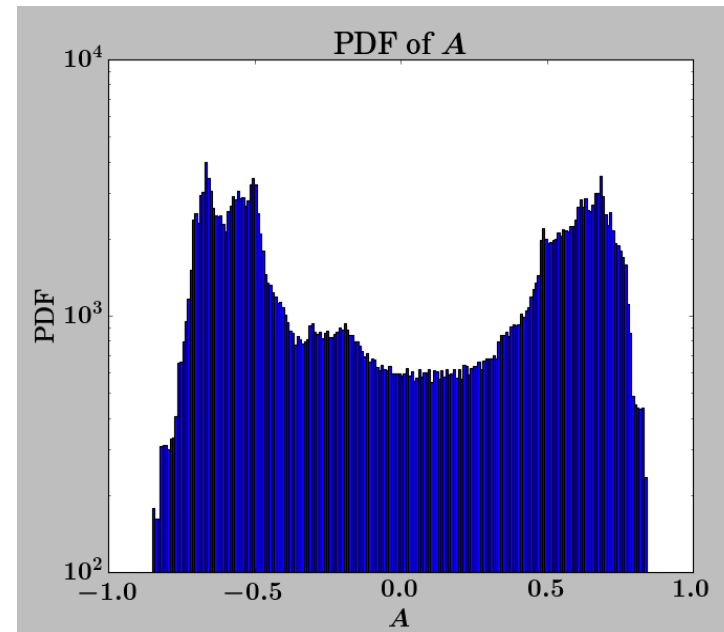
# Spatial Structure (Preliminary)

- Initial condition:  $\cos(x)$  for  $A$
- Shorter time (suppression phase)
  - Domains, and domain boundaries evident, resembles CHNS
  - A transport barriers?!
- Longer time (Ohmic decay phase)
  - Well mixed
  - No evidence nontrivial structure



# Something New, Cont'd

- For analysis: pdf of  $A$
- Suppression phase:
  - quenched diffusion
  - bi-modal distribution
    - quenching prevents fill-in
    - consequence i.c.
- Ohmic decay phase:
  - uni-modal distribution returns





# Conclusion, of Sorts

- Elastic fluids ubiquitous, interestingly similar and different. Comparison/contrast is useful approach.
- CHNS is interesting example of elastic turbulence where energy cascade is *not* fundamental or dominant.
- Spatio-temporal dynamics of (bi-stable) active scalar transport is a promising direction. Pattern formation in this system is terra novo.
- Revisiting polymer drag reduction would be interesting.