A Mean Field Model of the L \rightarrow H Transition in a Stochastic Magnetic Field W.X. Guo^{1*}, M. Jiang^{2**}, P.H. Diamond^{2,3}, C.-C. Chen³

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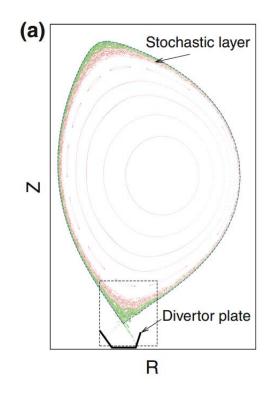
2020 AAPPS-DPP e-conference

Outline

- Motivation and Background
- Mean Field Model (a work in progress)
 - > What determines $\langle v'_E \rangle$? Key novelty: $\langle J_r \rangle$ induced by magnetic perturbation
 - $\rightarrow \langle V_{\theta} \rangle$ and $\langle V_{\phi} \rangle$ evolution
 - ightarrow Particle and heat flux
 - **Turbulence** intensity
 - > Modified 1D "Predator-Prey" model for $\langle E_r \rangle'$
- □ Lessons learned, so far
- Conclusion and Open Issues

Stochastic magnetic field

- The phenomenon of chaos of magnetic field lines is known as magnetic stochasticity (or magnetic chaos)
- In early studies, magnetic stochasticity is thought to be bad for confinement due to the enhanced radial transport of particles and energy along the chaotic field lines.
- However, at end of 1970s, magnetic stochasticity can be used to control the transport of energy and particles.
 → stochasticity as a positive: ergodic divertor concept



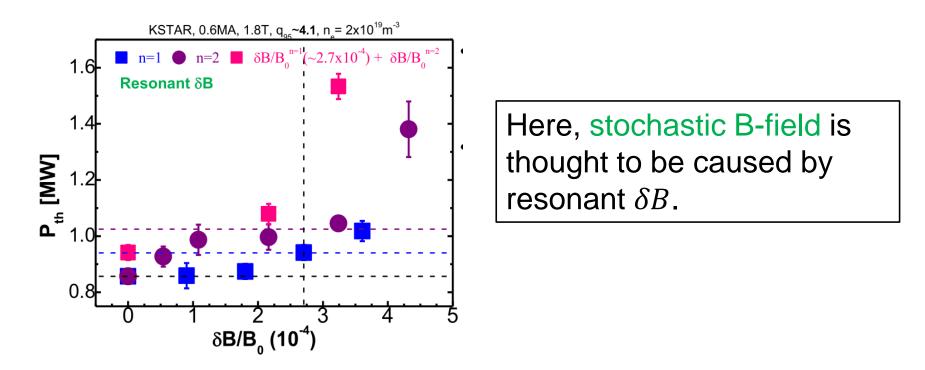
Typical Poincaré section of field lines in DIII-D like plasmas

Stochastic magnetic field induced by RMP

- Most of today's tokamaks need to use RMP (resonant magnetic perturbation) before L→H transition, in order to mitigate/suppress large ELMs, including first
- RMPs are thought to produce stochastic layer near separatrix
- The stochastic B field observed to influence
 - threshold and power
 - $-V_{E \times B}$ and flows
 - fluctuations

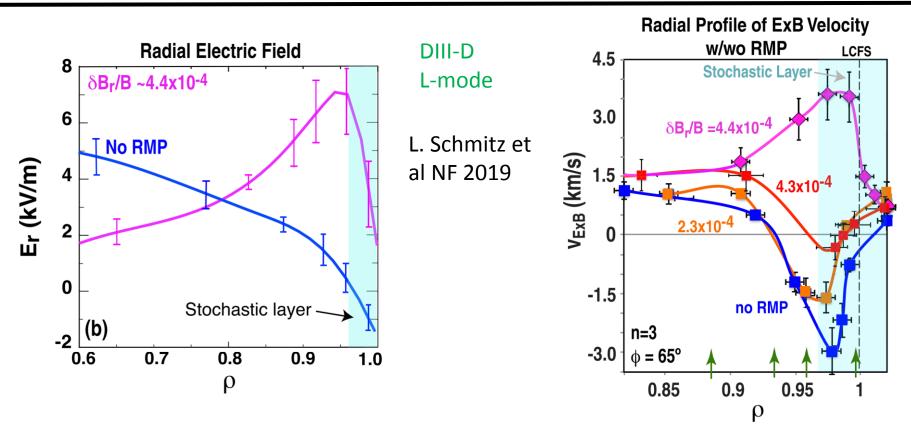
Motivation (why?)

RMPs increase P_{th} of L→H transition (experiments on DIII-D, MAST, AUG and KSTAR).



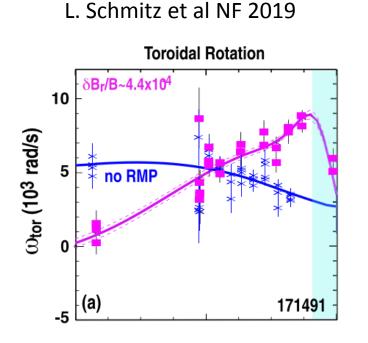
Need model to elucidate the physical mechanism.

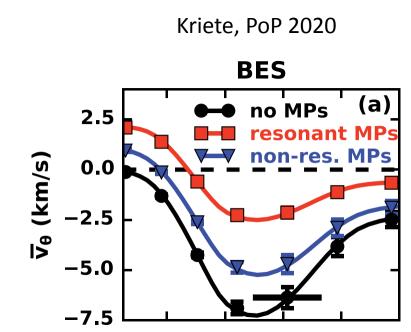
$\langle E_r \rangle$ structure: reversal by RMPs



- ✓ E_r reversal or 'bifurcation' : E_r well reduced or inverted to E_r hill with increasing RMP. Edge E_r shear layer sits in stochastic field region.
- ✓ Clear Change in E'_r due to increase of RMP field

Reduced toroidal/poloidal flow by RMPs





 RMP increased toroidal (co)rotation and shear at separatrix RMP reduces mean turbulence poloidal velocity



Clear effect on flows, and would further affect

$$E_r = \frac{1}{enB} \frac{\partial}{\partial r} P_{\rm i} + v_{\phi} B_{\theta} - v_{\theta} B_{\phi}$$

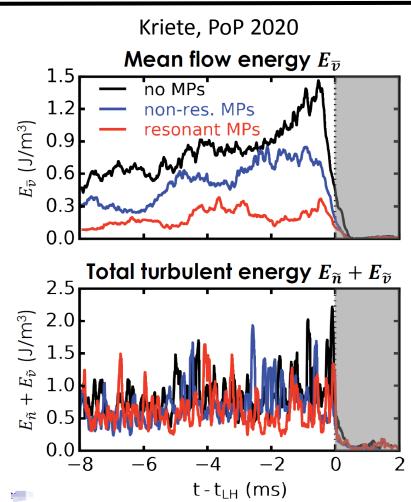
Fluctuations point: RMPs degrade energy transfer to mean flows prior to L-H transition

Energy balance between turbulence and mean flow^{1,2} Mean flow energy: $E_{\overline{v}} = \frac{1}{2}n_0m_i\langle\overline{v}_\theta\rangle^2$

Turbulent flow energy: $E_{\widetilde{v}} = \frac{1}{2}n_0m_i(\langle \widetilde{v}_r^2 \rangle + \langle \widetilde{v}_{\theta}^2 \rangle)$

Thermal free energy: $E_{\widetilde{n}} = \frac{1}{2}n_0T_{e0}(\widetilde{n}_e/n_0)^2$

- ✓ $E_{\bar{v}}$ increases prior to L-H in axisymmetric case
- ✓ $E_{\bar{v}}$ decreases due to RMPs, but total turbulent energy only changes slightly
- Power transfer to zonal flow decreases



Need more power to supply sufficient energy to the mean flow

Reduced Reynold stress/force by RMP

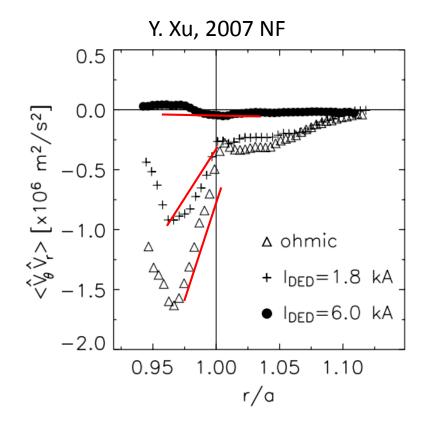
5.0

2.5

0.0

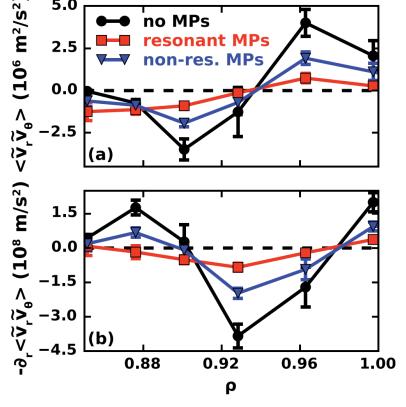
-2.5

(a)



✓ Increase I_{DED} reduces Reynold force.

Kriete, PoP 2020 sonant MPs MPs



✓ RMP reduces Reynold force in Lmode, thus increasing power required for H-mode.

Effects of RMP on transition? \leftrightarrow Model !

Model – $\langle E_r \rangle$ as critical

- **Goal:** ascertain change due to RMP mean field model for $\langle E_r \rangle$ and $\langle v'_E \rangle$
 - Radial force balance equation: $\langle E_r \rangle$ is determined by

- —Direct effect of $|\tilde{b}_r|^2$ on turbulence discussed later. Take turbulence as electrostatic in L-mode.
- —Prescribe stochastic B-field

$\langle J_r \rangle$ is key to L \rightarrow H trigger mechanism

• $L \rightarrow H$ transition:

 V'_E shearing feedback (mainly ∇P) \blacksquare "trigger" due to $\langle J_r \rangle$

• Several candidates for $\langle J_r \rangle$

—Radial flux of polarization charge \rightarrow turbulence intensity gradient scale

$$\rho_s^2 \langle \tilde{v}_r \nabla_{\!\!\perp}^2 \tilde{\phi} \rangle \rightarrow \frac{\partial}{\partial r} \langle \tilde{V}_\theta \tilde{V}_r \rangle$$
: Reynolds force

Critical Reynolds work criterion [Rich literatures: Diamond, Tynan et al]

- —neoclassical polarization $\rightarrow \rho_{\theta i}$ scale [S-I. Itoh, et al PRL, 60 (1988) 2276]
- —Orbit IOSS $\rightarrow \rho_{\theta i}$,..... [K.C.Shaing, PoF B: Plasma Physics 4 (1992) 171]

— New player here: $\langle J_r \rangle = \frac{\langle \tilde{J}_{\parallel} \tilde{B}_r \rangle}{B}$

stochastic field intensity profile scale, how enters?

Common element: $\langle J_r \rangle$ induced by stochastic B-field

• Ambipolarity breaking due to stochastic field $\Rightarrow \langle J_r \rangle$

$$\langle J_r \rangle = \left\langle \vec{J}_{\parallel} \cdot \vec{e}_r \right\rangle = \frac{\left\langle \tilde{J}_{\parallel} \widetilde{\mathbf{B}}_r \right\rangle}{B} \qquad \left\langle J_{\parallel} \right\rangle = \left\langle J_{\parallel,e} \right\rangle + \left\langle J_{\parallel,i} \right\rangle$$

• From Ampere law: $\tilde{J}_{\parallel} = -\frac{c}{4\pi} \nabla^2 \tilde{A}_{\parallel}$

Stochastic field produces currents in plasmas

$$\langle J_r \rangle = \frac{\langle \tilde{J}_{\parallel} \tilde{B}_r \rangle}{B} = -\frac{c}{4\pi B} \left\langle \frac{\partial}{\partial y} \tilde{A}_{\parallel} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) \tilde{A}_{\parallel} \right\rangle$$

$$= -\frac{c}{4\pi B} \frac{\partial}{\partial x} \left\langle \left(\frac{\partial}{\partial x} \tilde{A}_{\parallel} \right) \left(\frac{\partial}{\partial y} \tilde{A}_{\parallel} \right) \right\rangle = \frac{c}{4\pi B} \frac{\partial}{\partial x} \langle \tilde{B}_x \tilde{B}_y \rangle$$

$$= \frac{cB}{4\pi} \frac{\partial}{\partial x} \left\langle \tilde{b}_x \tilde{b}_y \right\rangle \Rightarrow \frac{cB_0}{4\pi} \frac{\partial}{\partial r} \left\langle \tilde{b}_r \tilde{b}_\theta \right\rangle \begin{bmatrix} \text{Maxwell} \\ \text{Force} \end{bmatrix}$$

Note: $\langle J_r \rangle$ tracks momentum, not heat transport.

Phase in Maxwell stress

- Maxwell stress $\langle \tilde{B}_r \tilde{B}_\theta \rangle = \sum_k |\tilde{A}_k|^2 \langle k_x k_y \rangle$
 - $\succ \tilde{A}_k$ tilted by developing $E \times B$ flow

$$\frac{\partial A}{\partial t} + V \cdot \nabla A = \eta J$$

$$\frac{\partial A}{\partial t} + V'_{E} \frac{\partial A}{\partial y} + \tilde{V} \cdot \nabla A = \eta J$$

$$k_{x} = k_{x}^{(0)} - k_{y} V'_{E} \tau_{c}$$

$$\tau_{c}: \begin{cases} shear \\ fluctuations \end{cases}$$

- > Thus, $\langle \tilde{B}_r \tilde{B}_\theta \rangle = -\sum_k |\tilde{A}_k|^2 \langle k_y^2 V'_E \tau'_c \rangle$
- Hence, Reynolds and Maxwell cross-phase closely linked.

 0	On How Decoherence of Vorticity Flux by Stochastic Magnetic Fields Quenches Zonal Flow Generation
San Diego	

Stochastic B-field affects $\langle V_{\theta} \rangle$

- Maxwell stress For V_{θ} : Turbulence of stochastic field Reynold stress perturbation Poloidal momentum balance $\frac{\partial \langle V_{\theta} \rangle}{\partial t} = -\mu(\langle V_{\theta} \rangle - V_{\theta,neo}) - \frac{\partial}{\partial r} \left(\left\langle \tilde{V}_{\theta} \tilde{V}_{r} \right\rangle - \frac{1}{4\pi o} \left\langle \tilde{B}_{r} \tilde{B}_{\theta} \right\rangle \right)$ For SS: $\langle V_{\theta} \rangle = V_{\theta,neo} - \frac{1}{\mu} \frac{\partial}{\partial r} \left(\left\langle \tilde{V}_{\theta} \tilde{V}_{r} \right\rangle - \frac{1}{4\pi\rho} \left\langle \tilde{B}_{r} \tilde{B}_{\theta} \right\rangle \right)$ $= V_{\theta,neo} - \frac{1}{\mu} \frac{\partial}{\partial r} \left(\frac{1}{B^2} \tau_c V'_E \frac{I}{1 + \alpha V'_E} - \frac{B^2}{4\pi\rho} \tau'_c V'_E |\tilde{b}_r|^2 \right)$ with $\mu = \mu_{00} (1 + \frac{\nu_{CX}}{\nu_{ii}}) \nu_{ii} q^2 R^2 \quad V_{\theta,neo} \approx -1.17 \frac{\partial T_i}{\partial r} \qquad \tau_c' = (\frac{k_{\theta}^2 V_E'^2 D_T}{2})^{-1/3}$
 - □ V'_E phasing via tilt tends to align turbulence and stochastic B-field, which counteracts the spin-up of ⟨V_θ⟩.
 □ ∂/∂r | b̃_r |² not only | b̃_r |², plays arole.

Stochastic B-field affects $\langle V_{\phi} \rangle$

• For
$$V_{\phi}$$
: $\frac{\partial \langle V_{\phi} \rangle}{\partial t} + \nabla \cdot \langle \tilde{V}_{r} \tilde{V}_{\phi} \rangle = \frac{1}{\rho c} \langle J_{r} \rangle B_{\theta} + S_{M}$
 $\langle \tilde{V}_{r} \tilde{V}_{\phi} \rangle = -\chi_{\phi} \frac{\partial}{\partial r} \langle V_{\phi} \rangle, \quad \chi_{\phi} = \chi_{T} = \frac{\rho_{s}^{2} C_{s}}{L_{T}}, \quad S_{M} = S_{a} \exp(-\frac{r^{2}}{2L_{M,dep}^{2}})$
Only consider diffusive term. G.B. Momentum source tracks heat (from core).
 $\Rightarrow \frac{\partial \langle V_{\phi} \rangle}{\partial t} = \frac{\partial}{\partial r} \left(\chi_{\phi} \frac{\partial}{\partial r} \langle V_{\phi} \rangle \right) + \frac{1}{4\pi\rho} \frac{B_{\theta}}{B} \frac{\partial}{\partial r} \langle \tilde{B}_{r} \tilde{B}_{\theta} \rangle + S_{M}$
• For SS: $\frac{\partial}{\partial r} \left(\chi_{\phi} \frac{\partial}{\partial r} \langle V_{\phi} \rangle \right) = -\frac{V_{T_{i}}^{2} B_{\theta}}{\beta} \frac{\partial}{\partial r} \langle \tilde{b}_{r} \tilde{b}_{\theta} \rangle - S_{M}$ Stochasticity edge toroidal velocity, shear
 $\Rightarrow \frac{\partial}{\partial r} \langle V_{\phi} \rangle |_{r_{sep}} = -\frac{1}{\chi_{\phi}} \int_{0}^{r_{sep}} S_{M} dr - \frac{V_{T_{i}}^{2} B_{\theta}}{\beta \chi_{\phi} B} \langle \tilde{b}_{r} \tilde{b}_{\theta} \rangle |_{r_{sep}}$ $B_{\theta} \langle J_{r} \rangle$
Integrated external torque with $\langle \tilde{b}_{r} \tilde{b}_{\theta} \rangle = k_{y}^{2} V'_{E} \tau'_{c} |\tilde{b}_{r}|^{2}$

Stochastic B-field affects electron density flux

• For electron density : $\frac{\partial n_e}{\partial t} + \frac{1}{r} \frac{\partial}{\partial r} (r\Gamma_e) = S_p$

with
$$\Gamma_e = -(D_{neo} + D_T) \frac{\partial n}{\partial r} + \Gamma_{e, stoch} \leftarrow D_{neo} = (m_e/m_i)^{1/2} \chi_{i,n}$$

•
$$D_T \sim b D_{GB}$$
 with b<1

eo

16

$$S_p = \Gamma_a \frac{a - r + d_a}{L_{dep}^2} \exp(-\frac{(a + d_a - r)^2}{2L_{dep}^2})$$

The stochastic field can induce particle flux $(n_e = n_i)$:

$$\Gamma_{e, stoch} = \frac{c}{4\pi e} \langle \tilde{b}_r \nabla_{\perp}^2 \tilde{A}_{\parallel} \rangle + n \langle \tilde{V}_{\parallel,i} \tilde{b}_r \rangle$$
with
$$\langle \tilde{b}_r \tilde{J}_{\parallel} \rangle \langle \tilde{b}_r \tilde{J}_{\parallel,i} \rangle \cdot \langle \tilde{b}_r \tilde{b}_{\theta} \rangle phasing$$

$$\langle \tilde{b}_r \tilde{b}_{\theta} \rangle = -\frac{c}{4\pi e} \frac{\partial}{\partial r} \langle \tilde{b}_r \tilde{b}_{\theta} \rangle = -n \frac{D_B}{\beta} \frac{\partial}{\partial r} \langle \tilde{b}_r \tilde{b}_{\theta} \rangle \text{ via } V'_E \text{ tilt.}$$

$$\langle \tilde{N}_E \tilde{V}_{\parallel,i} \tilde{b}_r \rangle: \text{ parallel ion flow along tilted field lines}$$

$$16$$

Stochastic B-field affects electron density flux

• The stochastic field can induce particle flux :

$$\Gamma_{stoch} = \frac{c}{4\pi e} \left\langle \tilde{b}_r \nabla_{\perp}^2 \tilde{A}_{\parallel} \right\rangle + n \left\langle \tilde{V}_{\parallel,i} \tilde{b}_r \right\rangle$$

• And from $\nabla_{\parallel}V_{\parallel} \cong 0$, $\tilde{v}_{\parallel,i} \sim -\tilde{b}_r \frac{\partial \langle V_{\parallel,i} \rangle}{\partial r} / (ik_{\parallel})$

$$\rightarrow \quad \left\langle \tilde{b}_{r} \tilde{V}_{\parallel,i} \right\rangle = -D_{M,\,eff} \, \frac{\partial \langle V_{\parallel,i} \rangle}{\partial r} = -\sum_{k} \langle \tilde{b}_{r}^{2} \rangle_{k} \, l_{ac,eff} \, \frac{\partial \langle V_{\parallel,i} \rangle}{\partial r}$$

Modify l_{ac} due to $E \times B$ shear

$$l_{ac,eff} = \frac{l_{ac}l_{V'_E}}{l_{ac}+l_{V'_E}}, \ l_{ac} = qR, l_{V'_E} = \frac{C_s}{|k_{\theta}\Delta V'_E|} \sim \frac{C_s}{|V'_E|}$$

✓ V_{pinch} is not diffusive, induced by stochastic B field
 ✓ Flow gradient drives particle flux.

Ion heat flux with stochastic field

For ion temperature : $n \frac{\partial T_i}{\partial t} + \frac{1}{r} \frac{\partial}{\partial r} (rQ_i) = S_H$

with Electrostatic

$$Q_{i} = -n(\chi_{i,neo} + \chi_{i,T})\nabla T_{i} + Q_{i,stoch} \quad \begin{cases} \cdot \chi_{i,neo} = \varepsilon^{-3/2}q^{2}\rho_{s}^{2}\nu_{ii} \\ \cdot \nu_{ii} = \frac{n_{0}Z^{4}e^{4}\ln\Lambda}{\sqrt{3}6\pi\varepsilon_{0}^{2}m_{i}^{1/2}T_{i0}^{3/2}} \\ \cdot \chi_{i,T} = \left(\frac{C_{s}^{2}\tau_{c}}{1+\alpha V_{E}^{\prime 2}}\right) * I \\ \cdot \chi_{GB} * I \end{cases}$$

The stochastic field affects ion heat flux :

$$Q_{i,stoch} = \int V_{\parallel} \langle \tilde{B}_r \delta f \rangle (V_{\parallel}^2 + V_{\perp}^2) = -\frac{\partial \langle T_i \rangle}{\partial r} \sqrt{\chi_{\parallel,i} \chi_{\perp,i}} \langle \tilde{b}_r^2 \rangle l_{ac} k_{\perp}$$

$$\frac{\partial \langle T_i \rangle}{\partial r} = -\frac{\partial \langle T_i \rangle}{\partial r} \sqrt{\chi_{\parallel,i} \chi_{\perp,i}} \langle \tilde{b}_r^2 \rangle l_{ac} k_{\perp}$$

 $= -v_{th,i} D_{M,eff} \frac{\partial \langle U_{i} \rangle}{\partial r} \simeq -D_{B,i} D_{M} \sqrt{k_{\perp}^2 \frac{\partial \langle U_{i} \rangle}{\partial r}}$ Potentially important as threshold power is directly related to heat flux. \checkmark Power uptake determines turbulence and Reynolds force \checkmark

Turbulence intensity

• For turbulence intensity :

$$\frac{\partial}{\partial t}I = \frac{\gamma_L}{(1+\alpha V_E'^2)^{\sigma}}I - \beta I^2 \to I = \frac{\gamma_L}{(1+\alpha V_E'^2)^{\sigma}}\frac{1}{\beta}$$

 $(\gamma_L - \text{growth rate}, \beta - \text{nonlinear decay rate}, V'_E - E \times B \text{ shear rate})$

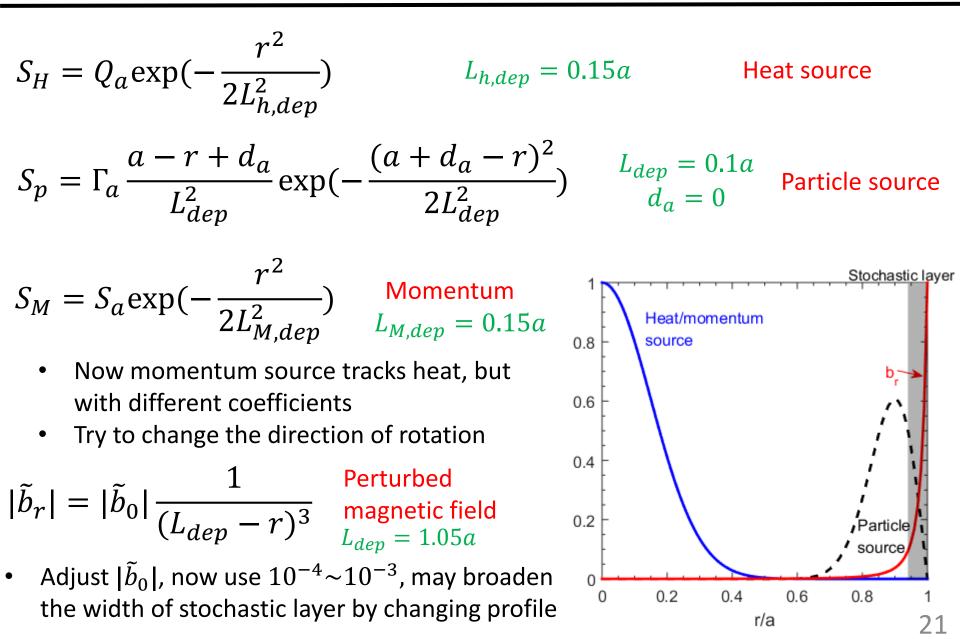
$$\frac{\gamma_L}{(1+\alpha V_E'^2)^{\sigma}} \sim \boldsymbol{\beta} \qquad \boldsymbol{\alpha} = \frac{1}{\alpha_0 (C_s/(qR))^2}, \quad \boldsymbol{\alpha}_0 \text{ and } \sigma \text{ are adjustable.}$$
$$\gamma_L = \gamma_{L0} \left(\frac{C_s}{R}\right) \sqrt{\frac{R}{L_T} - (\frac{R}{L_T})_{crit}} \sim \gamma_{L0} \left(\frac{C_s}{R}\right), \gamma_{L0} \approx 0.01$$
$$\text{threshold}$$

So, mean field "Predator-Prey" model

- RFB equation: $\langle V_E \rangle' = \frac{1}{eB} \frac{\partial}{\partial r} \langle \nabla P_i / n \rangle \frac{\partial}{\partial r} \langle V_\theta \rangle + \frac{B_\theta}{B} \frac{\partial}{\partial r} \langle V_\phi \rangle$
- Five field model:

 $\langle J_r \rangle$ directly related to Γ_e and Maxwell stress in flows

The sources



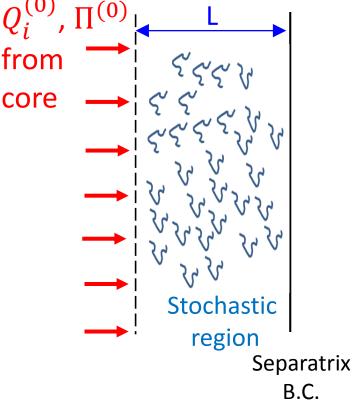
Lessons learned

• Novel $\langle J_r \rangle$ enters due to Stochastic field

- ✓ Generates Maxwell stress/force, which competes with Reynolds stress/force
- ✓ Effects on
 - -flow (poloidal and toroidal velocity)
 - phase
 - transport (particle and thermal)
- ✓ Reynolds and Maxwell phases linked by shearing

$$\left< \tilde{b}_r \tilde{b}_\theta \right> = k_y^2 V_E' \tau_c' |\tilde{b}_r|^2$$

Ongoing: toward to 1D solution



Reduced model as before, the model with:

 $- |\tilde{b}_r|^2$

Amplitude; scale

- Reduced RS in $\langle V_{\theta} \rangle$

Quenching the spin-up of V_{θ}

- Intrinsic torque in $\langle V_{\phi} \rangle$
- Effect of $|\tilde{b}_r|^2$ on Q_i , Γ_e

Goals:

> Output the time evolution of *I*, V_φ, V'_E, n, T_i for L→H physics modified by stochastic B-field.
 > Q_{i,crit}(| b̃_r|², ∂/∂r | b̃_r|²....)

Conclusions

- Ambipolarity breaking $\Rightarrow \langle \tilde{b}_r \tilde{b}_\theta \rangle$, contribute to $\langle J_r \rangle$
- Both amplitude and profile of $|b_r|^2$ matter.
- V'_E phasing \Rightarrow stochastic $\langle \tilde{b}_r \tilde{b}_\theta \rangle$ opposes turbulence $\langle \tilde{V}_r \tilde{V}_\theta \rangle$ phase linked
- Intrinsic toroidal torque, such that reversal or spin-up of edge $\langle V_{\phi} \rangle$ occur with RMP, $\langle \tilde{b}_r \tilde{b}_{\theta} \rangle$ enters edge $\langle V_{\phi} \rangle$
- $|b_r|^2$ can modify T_i and n_e profiles

.

Open issues

- Direct effect of stochasticity on turbulence $(l_K v.s. l_{\parallel} v.s. l_{mfp})$
- Collisionality scaling
- Cost of P_{th} to obtain H-mode with $E_r > 0$ (E_r hill) \Rightarrow Trade-off for particle control?
- RMP→island + stochastic region

↑ Mode length

Thank you very much !



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