

Ballooning mode in a Stochastic Magnetic Field —A Quasi-mode Model

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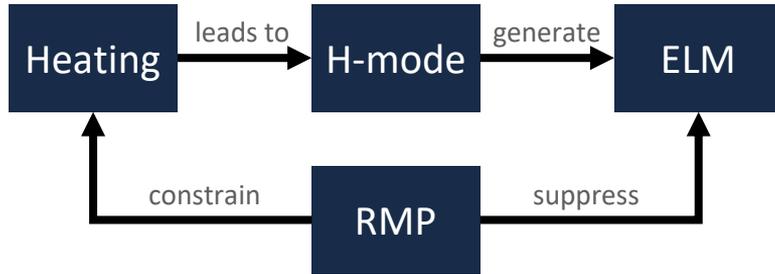
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- 2 Challenge: disparate geometries in the theories of ballooning mode and RMP
- 3 Model: quasi-mode, the counterpart of ballooning mode in a cylinder
- 4 Results: lessons learned for ballooning modes in a stochastic magnetic field
- 5 Future: suggested experimental and theoretical studies

Background: theoretical progress



Kobayashi, Masahiro, et al. *Physical review letters* 128.12 (2022): 125001.



Question: how does stochastic magnetic field modify plasma turbulence & instability process?

Instability and turbulent relaxation in a stochastic magnetic field

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The first model which

- maintains $\nabla \cdot \mathbf{J} = 0$ at all scales
- connects dynamics at micro and macro scales

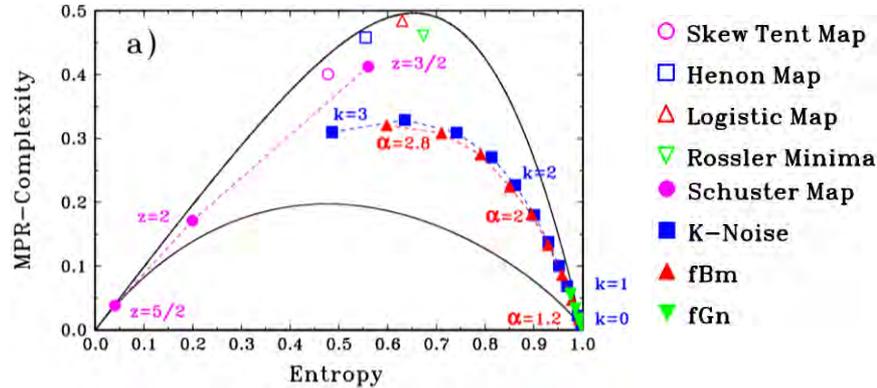
Experiments are also needed...

However, spectral analysis is not enough to characterize the turbulence state

Need other ways to study the statistics of plasma turbulence

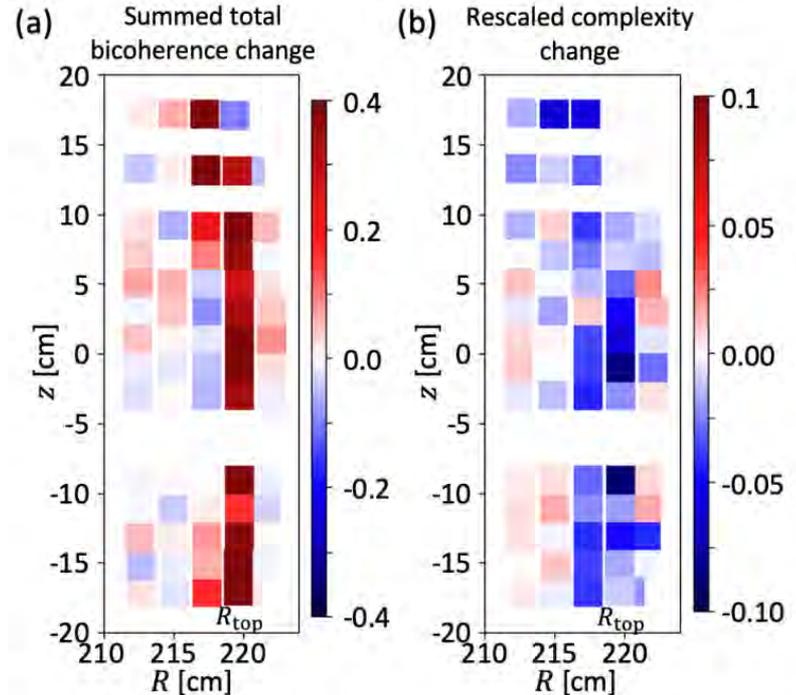
Background: recent experiment

- Recent experiments on KSTAR by Choi et al.^[1] → **complexity-entropy analysis**^[2]



- Jensen-Shannon complexity C_{JS}** : a metric of system's predictability. chaos: high C_{JS} ; noise: low C_{JS}
- In RMP ELM suppression phase, the rescaled complexity of edge temperature fluctuation reduces → **from "chaotic" turbulence to "noisy" turbulence**

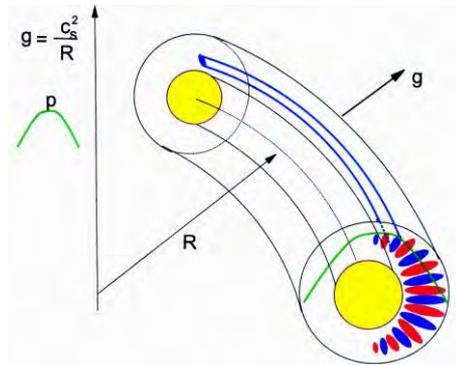
- Choi, Minjun J., et al. *Physics of Plasmas* 29.12 (2022).
- Rosso, Osvaldo A., et al. *Physical review letters* 99.15 (2007): 154102.



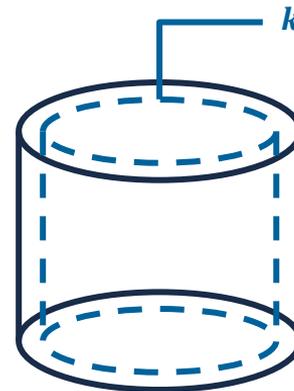
Explanation on the change in the turbulence statistical behavior?

Challenge: disparate geometries

- From interchange to a more **relevant** instability \longrightarrow ballooning mode
- A hard nut to crack: difference in geometries on which theories of the ballooning mode and RMP are based.



Ballooning mode
 \Updownarrow
 Toroidicity effect



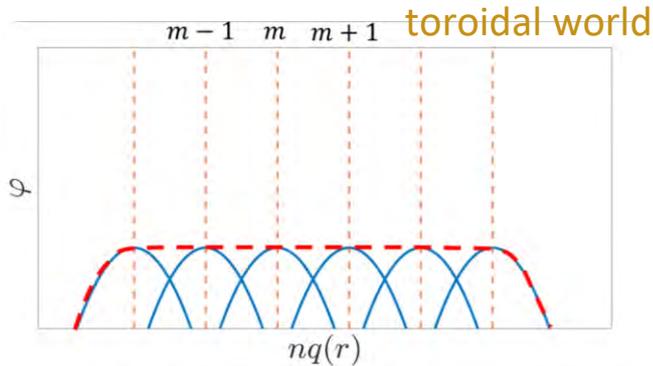
RMP
 \Updownarrow
 Resonant surfaces in a **cylinder**

Ballooning mode in a **torus** vs. resonant magnetic perturbations in a **cylinder**

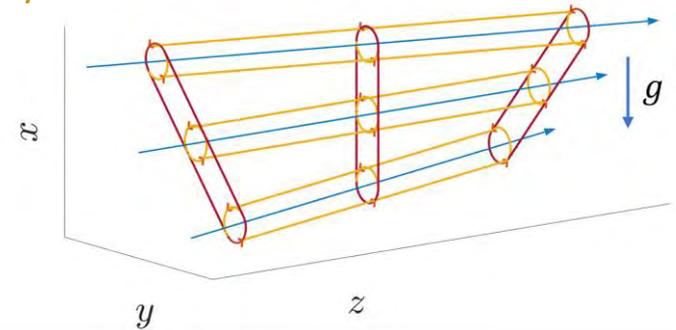
Question: Is there a way to circumvent this problem?

Strategy: find the counterpart of the ballooning mode

Theories of ballooning mode and RMP “reside” in different “parallel universe”.



cylindrical world



- Ballooning mode: a coupling of localized poloidal harmonics at different resonant surfaces

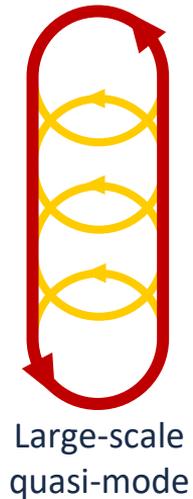
- Quasi-mode: a superposition of localized interchange modes at different resonant surfaces
- Broad in x , finite in z

Takeaway: a quasi-mode in a cylinder resembles a ballooning mode in a torus

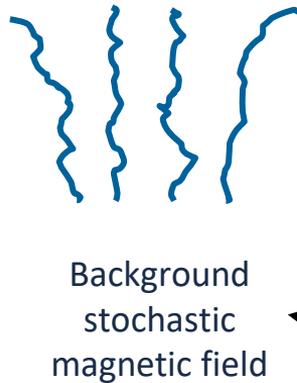
A Multi-scale Model: quasi-mode in a stochastic magnetic field

— Two-step scheme:

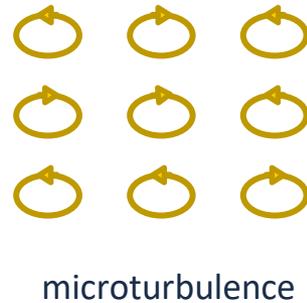
- Step 1: study the quasi-mode in a stochastic magnetic field
- Step 2: generalize the results to the ballooning mode



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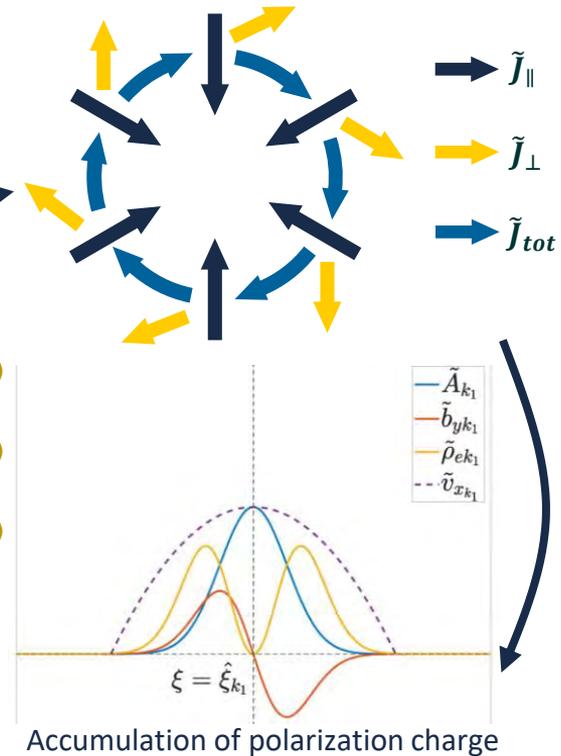


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Potential fluctuation $\tilde{\varphi}$

plasma could flow along chaotic field lines



A Multi-scale Model: quasi-mode in a stochastic magnetic field

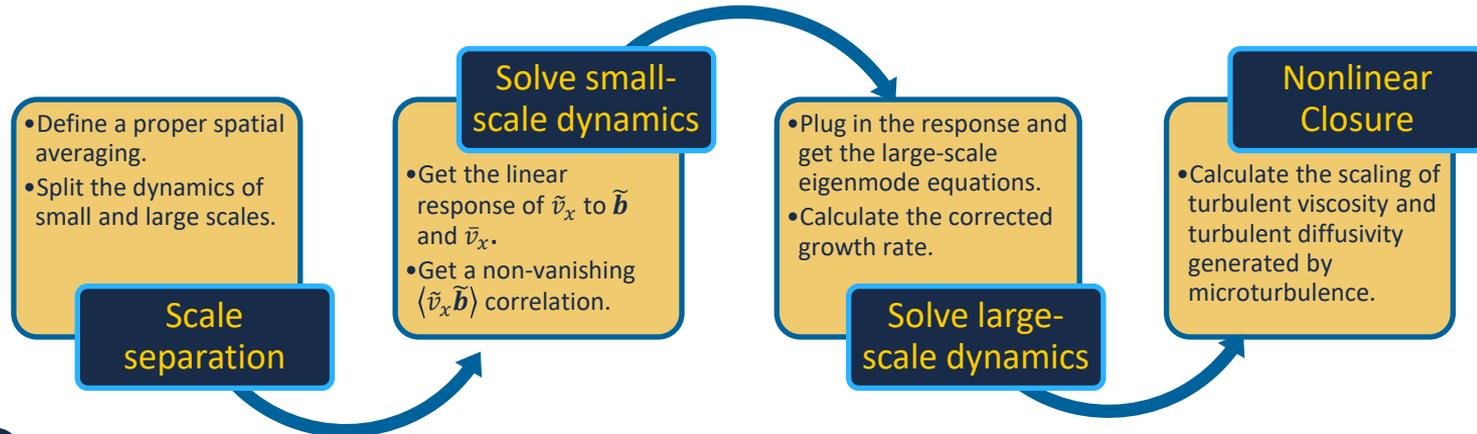
— The model of the quasi-mode is composed of vorticity equation and continuity equation

$$\left(\frac{\partial}{\partial t} - \underbrace{\nu_T \nabla_{\perp}^2}_{\text{turbulent viscosity}}\right) \nabla_{\perp}^2 (\bar{\varphi} + \tilde{\varphi}) + \frac{B_0^2}{\rho_0 \eta} \left(\frac{\partial}{\partial \zeta} + \tilde{\mathbf{b}} \cdot \nabla_{\perp}\right)^2 (\bar{\varphi} + \tilde{\varphi}) - \frac{g B_0}{\rho_0} \frac{\partial (\bar{\rho} + \tilde{\rho})}{\partial y} = 0 \quad \tilde{\varphi}, \tilde{\rho}, \tilde{v}_x: \text{microturbulence}$$

$$\left(\frac{\partial}{\partial t} - \underbrace{D_T \nabla_{\perp}^2}_{\text{turbulent diffusivity}}\right) (\bar{\rho} + \tilde{\rho}) = -(\bar{v}_x + \tilde{v}_x) \alpha \rho_0$$

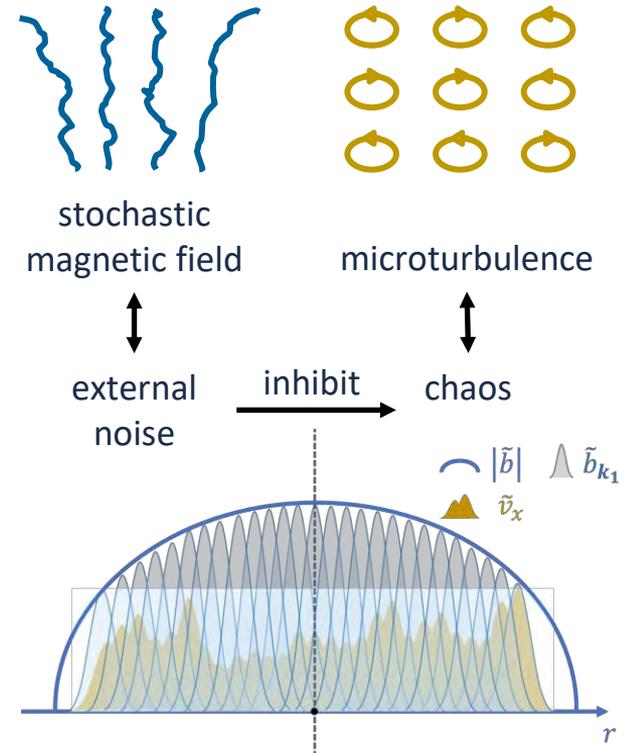
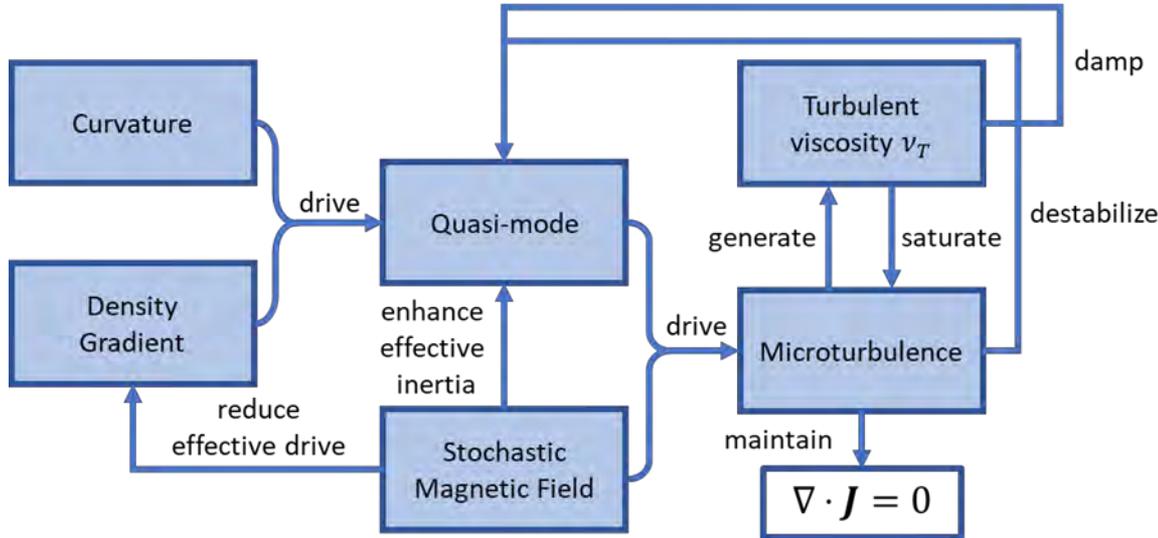
turbulent viscosity
turbulent diffusivity

comes from the
random advection of
the quasi-mode by
microturbulence



A Multi-scale Model: quasi-mode in a stochastic magnetic field

Results of this model can be summarized by a flowchart



Results: lessons we learned for ballooning mode

- microturbulence (small-scale convective cells) is driven $\longleftrightarrow \nabla \cdot J = 0$
 - increasing the number of triad interactions → **enhance nonlinear transfer**
 - **the increase in the bicoherence** observed in Choi's experiments

- Non-trivial correlation $\langle \tilde{b} \tilde{v}_x \rangle$ develops
 - Not only $\langle \tilde{b}_x \tilde{v}_x \rangle$, but also $\langle \tilde{b}_y \tilde{v}_x \rangle$ → absent in prior work
 - broad mode structure
 - the microturbulence **“locks on”** to the externally prescribed \tilde{b}
 - the edge plasma turbulence becomes **“noisy”**
 - the reduction in the C_{JS} in the RMP ELM suppression phase

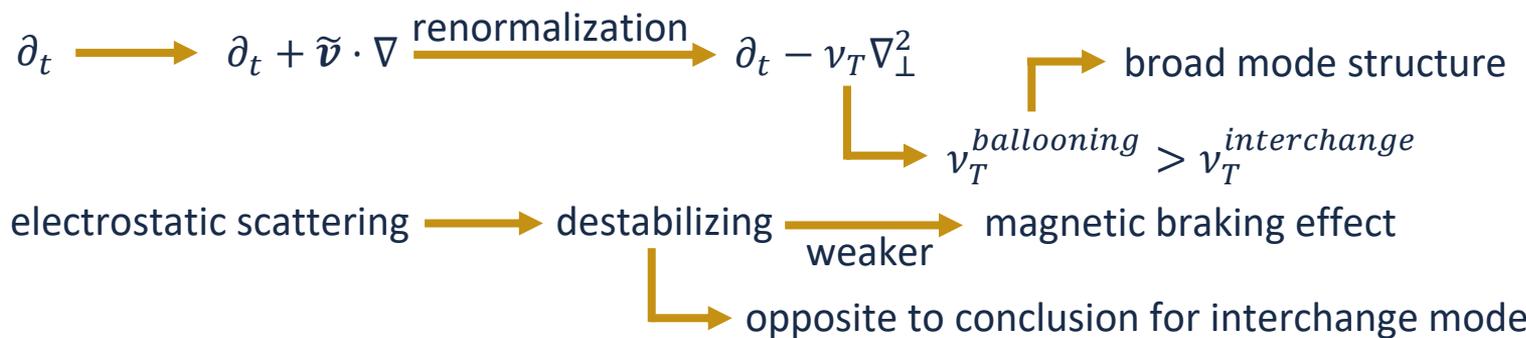
Results: lessons we learned for ballooning mode

— Stochastic magnetic field $\xrightarrow{\text{impede}}$ ballooning mode



1. Enhancing the effective plasma inertia (magnetic braking effect^[1])
2. Reducing the effective drive $\xrightarrow{\text{newly discovered}}$
3. Promoting turbulence damping

— Microturbulence $\xrightarrow{\text{drive}}$ **turbulent background**



1. Rutherford, Paul Harding. *The Physics of Fluids* 16.11 (1973): 1903-1908.

Future

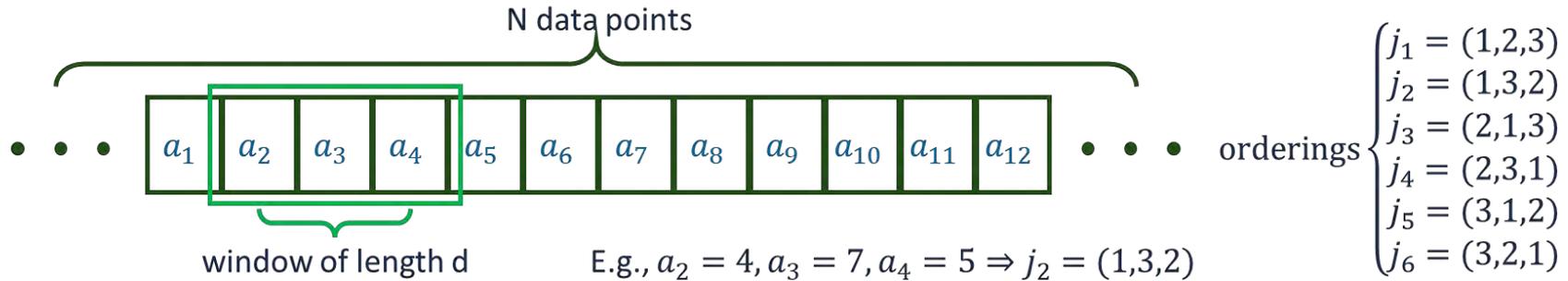
- Relate the reduction of the complexity to **dynamical** quantities? Suggested experiments:
 - i. Use **Beam Emission Spectroscopy** (BES) velocimetry to calculate the ratio of the **turbulent heat flux** to the **total heat flux**.
 - ii. Perform the **complexity-entropy analysis** for the velocity fluctuations collected from BES velocimetry during the RMP ELM mitigation or suppression phases
 - iii. Direct examination of the presence of the correlation $\langle \tilde{v}_x \tilde{\mathbf{b}} \rangle$.
- Directions for theoretical studies
 - i. Include **zonal flow** into our model. Hint: consider the magnetic shear.

$$dk_x/dt = k_x^{(0)} - \langle v_E \rangle' k_y, \quad dk_x/d_z = k_x^{(0)} - s k_y \rightarrow \langle k_x k_y \rangle \neq 0 / \langle \tilde{v}_x \tilde{v}_y \rangle \neq 0$$
 - ii. Study the **$Ku > 1$** case. Hint: **percolation theory**?
 - iii. Investigate the effect of stochastic magnetic field on blob propagation and SOL broadening. Hint: Theory of mean E×B shear in a stochastic magnetic field^[1]

1. Guo, Weixin, et al. *Plasma Physics and Controlled Fusion* 64.12 (2022): 124001.

Thank you!

Jensen-Shannon Complexity



Get the probability distribution function of orderings $P = \{p_j\}_{j=1,\dots,d!}$ ($N \gg d!$)

Then [1]

$$C_{JS} = \underbrace{H}_{\text{Shannon entropy}} \times \underbrace{Q}_{\text{J-S divergence}}$$

$$H = \frac{S}{S_{max}}, S = - \sum_j p_j \ln(p_j), S_{max} = \ln(d!)$$

$$Q = Q_0 \left\{ S \left(\frac{P + P_e}{2} \right) - \frac{S(P)}{2} - \frac{S(P_e)}{2} \right\}, P_e = \left\{ p_j = \frac{1}{d!} \right\}$$

1. Rosso, O.A., et al., *Physical review letters*, 99(15), p.154102.

